

Relaxation approximation of the Kerr Model for the three dimensional initial-boundary value problem

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Abstract. The electromagnetic wave propagation in a nonlinear medium is described by the Kerr model in the case of an instantaneous response of the material, or by the Kerr-Debye model if the material exhibits a finite response time. Both models are quasilinear hyperbolic and are endowed with a dissipative entropy. The initial-boundary value problem with a maximal-dissipative impedance boundary condition is considered here. When the response time is fixed, in both the one-dimensional and two-dimensional transverse electric cases, the global existence of smooth solutions for the Kerr-Debye system is established. When the response time tends to zero, the convergence of the Kerr-Debye model to the Kerr model is established in the general case, i.e. the Kerr model is the zero relaxation limit of the Kerr-Debye model.

AMS Classification. 35L50, 35Q60.

Key words. Initial-boundary value problem, Kerr model, Kerr-Debye model, relaxation, nonlinear Maxwell equations.

1 Introduction.

Nonlinear Maxwell equations are used for modelling nonlinear optical phenomena and electromagnetic wave propagation is described by

$$\begin{aligned}\partial_t D - \operatorname{curl} H &= 0, \\ \partial_t B + \operatorname{curl} E &= 0, \\ \operatorname{div} D &= \operatorname{div} B = 0.\end{aligned}$$

The field quantities E and H represent the electric and magnetic fields, respectively, while D and B are the electric and magnetic displacements. The constitutive relations are given by

$$\begin{aligned}B &= \mu_0 H, \\ D &= \varepsilon_0 E + P,\end{aligned}$$

where P is the polarization.

We consider here a homogeneous isotropic nonlinear medium (a crystal, for instance), so that the polarization is nonlinear. The Kerr model describes an instantaneous response of the medium and, in this case, P is given by

$$P = P_K = \varepsilon_0 \varepsilon_r |E|^2 E.$$

If the medium exhibits a finite response time τ , one should use the Kerr-Debye model for which

$$P = P_{KD} = \varepsilon_0 \chi E,$$

where

$$\partial_t \chi + \frac{1}{\tau} \chi = \frac{1}{\tau} \varepsilon_r |E|^2.$$

(For further details we refer the reader to, for example, [17] or [20].)

The Kerr-Debye model is a relaxation approximation of the Kerr model where τ is regarded as the relaxation parameter. (For a general discussion of relaxation problems, see [15].) Formally, when τ tends to 0, χ converges to $\varepsilon_r |E|^2$ and P_{KD} converges to P_K .

To cope with a physically realistic situation, we need to cover the initial-boundary value problem (IBVP), although in the rather simple geometry. We assume that the nonlinear material is confined in the region $\{x_1 \geq 0\}$, and we set $\Omega =]0, +\infty[\times \mathbb{R}^2$, $\Gamma = \{0\} \times \mathbb{R}^2$ denoting its boundary. We consider the Kerr and the Kerr-Debye models in the domain $\mathbb{R}_t^+ \times \Omega = [0, +\infty[\times \Omega$ with the impedance boundary condition on $\mathbb{R}_t^+ \times \Gamma$ and with zero initial data.

Once non-dimensionalized, the initial and boundary value problem (IBVP) for the Kerr model takes the following form (for $(t, x) \in \mathbb{R}^+ \times \Omega$):

$$\begin{aligned}\partial_t D - \operatorname{curl} H &= 0, \\ \partial_t H + \operatorname{curl} E &= 0,\end{aligned}$$

with the constitutive relation

$$D = (1 + |E|^2)E. \quad (1.1)$$

We suppose that the initial data vanish:

$$D(0, x) = H(0, x) = 0 \quad \text{for } x \in \Omega,$$

so that the divergence relations hold

$$\operatorname{div} D = \operatorname{div} H = 0 \quad \text{for } t \geq 0.$$

We denote by $n = {}^t(-1, 0, 0)$ the outer unit normal along Γ and we consider the impedance boundary condition

$$H \times n + a((E \times n) \times n) = \varphi \quad \text{for } (t, x) \in \mathbb{R}^+ \times \Gamma, \quad (1.2)$$

where a is a positive linear map on Γ . If $a = Id$, (1.2) is the classical ingoing wave condition. If $a = 0$, (1.2) is a Dirichlet boundary condition. The system is at rest for $t \leq 0$. It is only excited by the source term φ which is localized in the variable (t, x_2, x_3) and takes its values in Γ .

In the same fashion, the IBVP for the Kerr-Debye model (in which τ is replaced by ε) writes (for $(t, x) \in \mathbb{R}^+ \times \Omega$)

$$\begin{aligned}\partial_t D_\varepsilon - \operatorname{curl} H_\varepsilon &= 0, \\ \partial_t H_\varepsilon + \operatorname{curl} E_\varepsilon &= 0,\end{aligned} \quad (1.3)$$

$$\partial_t \chi_\varepsilon = \frac{1}{\varepsilon}(|E_\varepsilon|^2 - \chi_\varepsilon),$$

with the constitutive relation

$$D_\varepsilon = (1 + \chi_\varepsilon)E_\varepsilon. \quad (1.4)$$

We suppose that the initial data vanish:

$$D_\varepsilon(0, x) = H_\varepsilon(0, x) = 0, \quad \chi_\varepsilon(0, x) = 0 \quad \text{for } x \in \Omega,$$

and we have also

$$\operatorname{div} D_\varepsilon = \operatorname{div} H_\varepsilon = 0 \quad \text{for } t > 0.$$

In addition we suppose that we have the same impedance boundary condition

$$H_\varepsilon \times n + a((E_\varepsilon \times n) \times n) = \varphi \quad \text{for } (t, x) \in \mathbb{R}^+ \times \Gamma.$$

Two-dimensional models.

Following [19] we can also introduce the two-dimensional transverse magnetic (TM) and transverse electric (TE) models, as follows. For the transverse magnetic case we assume that

$$\begin{aligned}H(x_1, x_2, x_3) &= {}^t(0, H_2(x_1, x_3), 0), \\ E(x_1, x_2, x_3) &= {}^t(E_1(x_1, x_3), 0, E_3(x_1, x_3)),\end{aligned}$$

in the domain $(x_1, x_3) \in \{x_1 > 0\} \times \mathbb{R}$. The Maxwell system becomes

$$\begin{cases} \partial_t D_1 + \partial_3 H_2 = 0, \\ \partial_t D_3 - \partial_1 H_2 = 0, \\ \partial_t H_2 + \partial_3 E_1 - \partial_1 E_3 = 0, \end{cases}$$

while the divergence-free condition reads

$$\partial_1 D_1 + \partial_3 D_3 = 0.$$

(In this case the divergence condition for H is irrelevant.) The impedance boundary condition reads

$$H_2 - aE_3 = \varphi \text{ with } a \geq 0.$$

The above system is coupled with either (1.1) (in the Kerr model), or (1.4) and the third equation in (1.3) (in the Kerr-Debye model).

In the transverse electric case, we assume that

$$\begin{aligned} E(x_1, x_2, x_3) &= {}^t(0, E_2(x_1, x_3), 0), \\ H(x_1, x_2, x_3) &= {}^t(H_1(x_1, x_3), 0, H_3(x_1, x_3)), \end{aligned}$$

and we obtain

$$\begin{cases} \partial_t D_2 - \partial_3 H_1 + \partial_1 H_3 = 0, \\ \partial_t H_1 - \partial_3 E_2 = 0, \\ \partial_t H_3 + \partial_1 E_2 = 0, \end{cases} \quad (1.5)$$

with the divergence-free condition

$$\partial_1 H_1 + \partial_3 H_3 = 0. \quad (1.6)$$

The impedance boundary condition becomes

$$H_3 + aE_2 = \varphi \text{ with } a \geq 0. \quad (1.7)$$

For the Kerr model, (1.5)-(1.7) is coupled with

$$D_2 = (1 + (E_2)^2),$$

and for the Kerr-Debye model, (1.5)-(1.7) is coupled with

$$\begin{aligned} D_{2,\varepsilon} &= (1 + \chi_\varepsilon)E_{2,\varepsilon}, \\ \partial_t \chi_\varepsilon &= \frac{1}{\varepsilon}((E_{2,\varepsilon})^2 - \chi_\varepsilon). \end{aligned}$$

In the case of a fixed finite response time, numerical simulations are presented for these two-dimensional models by finite-difference methods in [19] and by finite-element methods in [13].

One-dimensional model.

In [3], the following one-dimensional model is introduced:

$$\begin{aligned} E(x_1, x_2, x_3) &= {}^t(0, e(x_1), 0), \\ H(x_1, x_2, x_3) &= {}^t(0, 0, h(x_1)). \end{aligned}$$

In this case, the IBVP for the Kerr model reads

$$\begin{aligned} \partial_t d + \partial_1 h &= 0, \\ \partial_t h + \partial_1 e &= 0, \\ d &= (1 + e^2)e, \end{aligned} \quad (1.8)$$

with the impedance boundary condition

$$h(t, 0) + ae(t, 0) = \varphi(t), \quad t \geq 0, \text{ where } a \geq 0, \quad (1.9)$$

and the zero initial data

$$e(0, x_1) = h(0, x_1) = 0, \quad x_1 \geq 0. \quad (1.10)$$

The one-dimensional IBVP for the Kerr-Debye model is

$$\begin{aligned} \partial_t d_\varepsilon + \partial_1 h_\varepsilon &= 0, \\ \partial_t h_\varepsilon + \partial_1 e_\varepsilon &= 0, \\ \partial_t \chi_\varepsilon &= \frac{1}{\varepsilon} ((e_\varepsilon)^2 - \chi_\varepsilon), \\ d_\varepsilon &= (1 + \chi_\varepsilon) e_\varepsilon, \end{aligned} \quad (1.11)$$

with the impedance boundary condition

$$h_\varepsilon(t, 0) + ae_\varepsilon(t, 0) = \varphi(t), \quad t \geq 0, \text{ where } a \geq 0, \quad (1.12)$$

and the null initial data

$$e_\varepsilon(0, x_1) = h_\varepsilon(0, x_1) = \chi_\varepsilon(0, x_1) = 0, \quad x_1 \geq 0. \quad (1.13)$$

We can also remark that the divergence conditions on h and d are irrelevant for both models.

Mathematical properties and main results.

Since both Kerr and Kerr-Debye models are endowed with strictly convex entropies, the associated systems are symmetrizable hyperbolic. Furthermore the boundary is characteristic of constant multiplicity (except for the system (1.8)) and the boundary conditions are maximal dissipative. So the general results in [10] ensure the local existence of smooth solutions for smooth data. For the Kerr model a better local existence result is proved in [16]. A similar result for the Kerr-Debye model is established in the Appendix.

We denote by T^* and T_ε^* the lifespan of such smooth solutions for the Kerr and the Kerr-Debye IBVP, respectively. Since the Kerr model is a homogeneous quasilinear hyperbolic system, shock waves can appear at T^* . In particular, for the one-dimensional case, we can rewrite (1.8) as a p-system which is genuinely nonlinear for $d \neq 0$. In this case, using the results of [14], we can exhibit initial data such that the lifespan T^* is finite with formation of shock waves (see also [4]). On the other hand the Kerr-Debye model is a quasilinear hyperbolic system with source term and it is totally linearly degenerate, i.e. each characteristic field is linearly degenerate. So we can expect that, if the lifespan T_ε^* is finite, the behavior of the smooth solution is analogous to the semilinear case. Indeed we obtained this result in the 1-d case in [5]: if T_ε^* is finite then the solution and its gradient explode, i.e. shock formation never occurs. In fact, using stronger dissipative properties for the Kerr-Debye model we proved in [7] that $T_\varepsilon^* = +\infty$ for the one-dimensional Cauchy problem. This global existence result should be also established for the Cauchy problem in the 2-d TE case. In the present paper we choose to prove the same result for the impedance IBVP in the 1-d and 2-d TE cases (see Theorem 1 below). In the proofs we must take into account a new difficulty which does not appear for the Cauchy problem: the boundary Γ of the domain Ω is characteristic.

The Kerr-Debye model is a relaxation approximation of the Kerr model in the sense developed in [8]. The stability conditions in [8] and [18] are satisfied so it is natural to study the behavior of the smooth solutions to the Kerr-Debye model as the relaxation coefficient ε tends to zero. Concerning the Cauchy problem with initial data (D_0, H_0, χ_0) satisfying $\operatorname{div} D_0 = \operatorname{div} H_0 = 0$, $\chi_0 \geq 0$, the convergence for the smooth solutions is proved in [11] using the results of [18]. Generally a boundary layer in time appears because of the non compatibility of the initial data with the equilibrium condition $\chi = |E|^2$. We study here the convergence for the impedance IBVP. In this case no boundary layer appears, neither in the time variable (since the zero initial data fit the equilibrium condition) nor in the space variables (since the boundary condition is the same for both IBVP). In the one-dimensional case, we presented in [3] a first convergence result. In fact we obtained the convergence of (1.11)-(1.12)-(1.13) to (1.8)-(1.9)-(1.10) on some interval $[0, \tilde{T}] \subset [0, T^*]$. The

same kind of convergence result is announced for the 3-d case in [6]. Here we improve these results in all cases since the convergence is obtained on each interval $[0, T] \subset [0, T^*[$.

Our paper is organized as follows. In Section 2 we exhibit general properties of the Kerr and the Kerr-Debye models. For a fixed ε we establish the global existence result of the solutions to the Kerr Debye 2-d TE model. Section 3 is devoted to convergence results when the relaxation parameter tends to zero.

2 General Properties.

2.1 Properties of the Kerr model.

We recall the initial-boundary value problem for the general Kerr model:

$$\begin{cases} \partial_t D - \operatorname{curl} H = 0, \\ \partial_t H + \operatorname{curl} E = 0, \\ D = (1 + |E|^2)E, \end{cases} \quad (2.1)$$

for $(t, x) \in \mathbb{R}^+ \times \Omega$ together with the initial and boundary conditions

$$D(t=0) = H(t=0) = 0 \text{ for } x \in \Omega, \quad (2.2)$$

$$H \times n + a((E \times n) \times n) = \varphi \text{ for } (t, x) \in \mathbb{R}^+ \times \Gamma. \quad (2.3)$$

The energy density given by

$$\mathcal{E}_K(D, H) = \frac{1}{2}(|E|^2 + |H|^2 + \frac{3}{2}|E|^4)$$

is a strictly convex entropy (with associated flux function $E \times H$), so (2.1) is a quasilinear hyperbolic symmetrizable system.

In the three-dimensional case, the eigenvalues are (for $\xi \neq 0$)

$$\lambda_1(E, \xi) \leq \lambda_2(E, \xi) < \lambda_3 = \lambda_4 = 0 < \lambda_5 = -\lambda_2 \leq \lambda_6 = -\lambda_1,$$

so the boundary $\mathbb{R}^+ \times \Gamma$ is characteristic of constant multiplicity two. By direct calculations we obtain

$$\lambda_1(E, \xi) = -(1 + |E|^2)^{-\frac{1}{2}}|\xi|,$$

$$\lambda_2(E, \xi) = -(1 + |E|^2)^{-\frac{1}{2}}(1 + 3|E|^2)^{-\frac{1}{2}}((1 + |E|^2)|\xi|^2 + 2(E \cdot \xi)^2)^{\frac{1}{2}}.$$

In the two-dimensional cases, TM and TE, the eigenvalues are of the form:

$$\lambda_1(E, \xi) < \lambda_2 = 0 < \lambda_3 = -\lambda_1,$$

so the boundary $\mathbb{R}^+ \times \Gamma$ is characteristic of constant multiplicity one. By direct calculations we obtain for the TM model:

$$\lambda_1^{TM}(E, \xi) = -(1 + |E|^2)^{-\frac{1}{2}}(1 + 3|E|^2)^{-\frac{1}{2}}((1 + |E|^2)|\xi|^2 + 2(E \cdot \xi)^2)^{\frac{1}{2}},$$

and for the TE model:

$$\lambda_1^{TE}(E, \xi) = -(1 + 3|E|^2)^{-\frac{1}{2}}|\xi|.$$

In the one-dimensional case, the system is strictly hyperbolic and the boundary is non characteristic. We have

$$\lambda_1(E) < 0 < \lambda_2 = -\lambda_1,$$

with

$$\lambda_1(E) = -(1 + 3e^2)^{-\frac{1}{2}}.$$

We remark that the impedance boundary condition (2.3) is maximal dissipative. Generally speaking local existence results of smooth solutions to IBVP for quasilinear hyperbolic systems with characteristic boundary are available in [10]. For the Kerr system (2.1)-(2.2)-(2.3) we can also apply a more adapted result in [16]. We assume that the source term φ is compactly supported in $\mathbb{R}^+ \times \Gamma$. We denote by H^s the classical Sobolev spaces and we suppose that φ belongs to $H^s(\mathbb{R}_t \times \Gamma)$ for s great enough. So the boundary condition (2.3) and the initial data (2.2) are compatible and by [16] we obtain smooth local solutions.

Proposition 1 *Let $\varphi \in H^3(\mathbb{R} \times \Gamma)$ compactly supported in $\mathbb{R}^+ \times \Gamma$. Then there exists a maximal smooth solution (D, H) to the IBVP (2.1)-(2.2)-(2.3) which lifespan is denoted by T^* and such that*

$$\partial_t^i(D, H) \in \mathcal{C}^0([0, T^*]; H^{3-i}(\Omega)) \text{ for } i = 0, 1, 2, 3.$$

Remark 1 *In the third part we need more regular solutions. In fact, using [10] we obtain the following result: for all $k \in \mathbb{N}$, there exists $s_0(k)$ such that if $\varphi \in H^{s_0(k)}(\mathbb{R} \times \Gamma)$ then*

$$\partial_t^i(D, H) \in \mathcal{C}^0([0, T^*]; H^{k-i}(\Omega)) \text{ for } i = 0, \dots, k.$$

For example, for $k = 7$, using theorem 2 in [10] we obtain that $s_0(7) = 18$.

The previous existence result is stated in the three-dimensional case. Analogous results can be obtained for both the two-dimensional and the one-dimensional cases.

2.2 Properties of the Kerr-Debye models.

Consider for a fixed $\varepsilon > 0$ the Kerr-Debye model

$$\begin{cases} \partial_t D_\varepsilon - \operatorname{curl} H_\varepsilon = 0, \\ \partial_t H_\varepsilon + \operatorname{curl} E_\varepsilon = 0, \\ \partial_t \chi_\varepsilon = \frac{1}{\varepsilon} (|E_\varepsilon|^2 - \chi_\varepsilon), \\ D_\varepsilon = (1 + \chi_\varepsilon) E_\varepsilon. \end{cases} \quad (2.4)$$

for $(t, x) \in \mathbb{R}^+ \times \Omega$ together with the initial and boundary conditions

$$D_\varepsilon(0, x) = H_\varepsilon(0, x) = 0, \chi_\varepsilon(0, x) = 0 \text{ for } x \in \Omega, \quad (2.5)$$

$$H_\varepsilon \times n + a((E_\varepsilon \times n) \times n) = \varphi \text{ for } (t, x) \in \mathbb{R}^+ \times \Gamma. \quad (2.6)$$

The divergence free conditions are preserved by the system:

$$\operatorname{div} H_\varepsilon = \operatorname{div} (1 + \chi_\varepsilon) E_\varepsilon = 0. \quad (2.7)$$

By the third equation in (2.4) we observe that we have

$$\chi_\varepsilon \geq 0.$$

The energy density given by

$$\mathcal{E}_{KD}(D, H, \chi) = \frac{1}{2}(1 + \chi)^{-1}|D|^2 + \frac{1}{2}|H|^2 + \frac{1}{4}\chi^2 \quad (2.8)$$

is a strictly convex entropy in the domain $\{\chi \geq 0\}$ (with associated flux function $E \times H = (1 + \chi)^{-1}D \times H$). So (2.4) is a quasilinear symmetrizable hyperbolic system.

In the three-dimensional case the eigenvalues are, for $\xi \neq 0$,

$$\mu_1(\chi, \xi) = \mu_2 < \mu_3 = \mu_4 = \mu_5 = 0 < \mu_6 = \mu_7 = -\mu_1,$$

where $\mu_1 = -(1 + \chi)^{-\frac{1}{2}}|\xi|$. So the boundary is characteristic of constant multiplicity three.

In the two-dimensional cases we obtain

$$\mu_1(\chi, \xi) < \mu_2 = \mu_3 = 0 < \mu_4 = -\mu_1,$$

where $\mu_1 = -(1 + \chi)^{-\frac{1}{2}}|\xi|$. So the boundary is characteristic of constant multiplicity two.

In the one-dimensional case, the system is strictly hyperbolic and the boundary is characteristic of constant multiplicity one. The eigenvalues are

$$\mu_1(\chi) < \mu_2 = 0 < \mu_3 = -\mu_1,$$

where $\mu_1 = -(1 + \chi)^{-\frac{1}{2}}$.

We remark that each characteristic field of the quasilinear hyperbolic system (2.4) is linearly degenerate. It suffices to prove this property for the eigenvalue $\mu_1(\chi, \xi) = -(1 + \chi)^{-\frac{1}{2}}|\xi|$ for which the last component of the corresponding eigenvector vanishes, so we have

$$\nabla \mu_1 \cdot r_1 \equiv 0.$$

In the 3-d case this property is also obtained by a general result in [2].

Using the special structure of the Kerr-Debye model, we prove in the Appendix the following existence result.

Proposition 2 *Let $\varphi \in H^5(\mathbb{R} \times \Gamma)$ compactly supported in $\mathbb{R}^+ \times \Gamma$. Then there exists a maximal smooth solution $(D_\varepsilon, H_\varepsilon, \chi_\varepsilon)$ to the IBVP (2.4)-(2.5)-(2.6) which lifespan is denoted by T_ε^* and such that*

$$\partial_t^i(D_\varepsilon, H_\varepsilon, \chi_\varepsilon) \in \mathcal{C}^0([0, T_\varepsilon^*]; H^{3-i}(\Omega)) \text{ for } i = 0, 1, 2, 3.$$

Analogous results can be established in both the 2-d and the 1-d cases.

2.3 Dissipative properties of the Kerr-Debye models.

For the Cauchy problem, it is well known (see [11]) that the Kerr-Debye system is dissipative and we have

$$\frac{d}{dt} \int_{\mathbb{R}^3} \mathcal{E}_{KD}(D, H, \chi) dx = -\frac{\varepsilon}{2} \int_{\mathbb{R}^3} |\partial_t \chi|^2 dx.$$

We generalize this dissipative formula for the IBVP (2.4)-(2.5)-(2.6). In addition we exhibit another dissipative property for the time derivatives. To start with we extend the boundary condition (2.6) in the following way: we replace H_2 by $H_2 + \varphi_3(t, x_2, x_3)\eta(x_1)$, H_3 by $H_3 - \varphi_2(t, x_2, x_3)\eta(x_1)$, where η is a cut off function which is equal to 1 in a neighbourhood of 0, compactly supported in the interval $[0, 1]$. Furthermore, to preserve the divergence free condition (2.7) we replace H_1 by

$$H_1 + (\partial_2 \varphi_3 - \partial_3 \varphi_2)(t, x_2, x_3) \int_{x_1}^{+\infty} \eta(s) ds$$

and we denote

$$\mathcal{R}(t, x) = \begin{pmatrix} (\partial_2 \varphi_3 - \partial_3 \varphi_2)(t, x_2, x_3) \int_{x_1}^{+\infty} \eta(s) ds \\ \varphi_3(t, x_2, x_3)\eta(x_1) \\ -\varphi_2(t, x_2, x_3)\eta(x_1) \end{pmatrix}.$$

Then the system (2.4)-(2.7) becomes in the variable $V = (U, \chi) = (E, H, \chi)$:

$$\begin{aligned} (i) \quad & (1 + \chi)\partial_t E + (\partial_t \chi)E - \text{curl } H = G_1, \\ (ii) \quad & \partial_t H + \text{curl } E = G_2, \\ (iii) \quad & \partial_t \chi = \frac{1}{\varepsilon}(|E|^2 - \chi), \end{aligned} \tag{2.9}$$

where $G_1 = \text{curl } \mathcal{R}$ and $G_2 = -\partial_t \mathcal{R}$, with the homogeneous initial and boundary conditions:

$$V(0, x) = 0, x \in \Omega, \quad (2.10)$$

$$H \times n + a((E \times n) \times n) = 0, (t, x) \in \mathbb{R}^+ \times \Gamma. \quad (2.11)$$

The divergence free conditions are preserved:

$$\text{div } H = \text{div } ((1 + \chi)E) = 0. \quad (2.12)$$

For the IBVP (2.9)-(2.12) the dissipation properties are described by the following result:

Proposition 3 *Let V be the smooth solution to (2.9)-(2.12) given by Proposition 2. We denote $E_T = (0, E_2, E_3)$. Then, on the interval $[0, T_\varepsilon^*[,$ we have*

$$\frac{d}{dt} \int_{\Omega} \mathcal{E}_{KD}(D, H, \chi) dx + \frac{\varepsilon}{2} \int_{\Omega} |\partial_t \chi|^2 dx + \int_{\Gamma} a(E_T) \cdot E_T dx_2 dx_3 = \int_{\Omega} (G_1 \cdot E + G_2 \cdot H) dx, \quad (2.13)$$

$$\begin{aligned} & \frac{1}{2} \frac{d}{dt} \int_{\Omega} \left((1 + \chi) |\partial_t E|^2 + |\partial_t H|^2 + \frac{1}{2} (\partial_t \chi)^2 \right) dx + \frac{3}{2\varepsilon} \int_{\Omega} |E|^2 |\partial_t E|^2 dx + \frac{\varepsilon}{2} \int_{\Omega} |\partial_{tt} \chi|^2 dx \\ & + \int_{\Gamma} a(\partial_t E_T) \cdot \partial_t E_T dx_2 dx_3 = \frac{3}{2\varepsilon} \int_{\Omega} \chi |\partial_t E|^2 dx + \int_{\Omega} (\partial_t G_1 \cdot \partial_t E + \partial_t G_2 \cdot \partial_t H) dx. \end{aligned} \quad (2.14)$$

Proof. Taking the inner product of (2.9.i) with E and (2.9.ii) with H we obtain

$$\int_{\Omega} (1 + \chi) \partial_t E \cdot E dx + \int_{\Omega} |E|^2 \partial_t \chi dx - \int_{\Omega} \text{curl } H \cdot E dx + \int_{\Omega} \partial_t H \cdot H dx + \int_{\Omega} \text{curl } E \cdot H dx = \int_{\Omega} (G_1 \cdot E + G_2 \cdot H) dx. \quad (2.15)$$

After integrating by parts we obtain

$$\begin{aligned} - \int_{\Omega} \text{curl } H \cdot E dx + \int_{\Omega} \text{curl } E \cdot H dx &= \int_{\Gamma} (H \times n) \cdot E dx_2 dx_3 \\ &= - \int_{\Gamma} a((E \times n) \times n) \cdot E dx_2 dx_3 \\ &= \int_{\Gamma} a(E_T) \cdot E_T dx_2 dx_3. \end{aligned}$$

In addition, we have

$$\int_{\Omega} (1 + \chi) \partial_t E \cdot E dx = \frac{1}{2} \frac{d}{dt} \int_{\Omega} (1 + \chi) |E|^2 dx - \frac{1}{2} \int_{\Omega} |E|^2 \partial_t \chi dx.$$

Using the last equation in (2.9), we obtain

$$\frac{1}{2} \int_{\Omega} |E|^2 \partial_t \chi dx = \frac{1}{4} \frac{d}{dt} \int_{\Omega} |\chi|^2 dx + \frac{\varepsilon}{2} \int_{\Omega} (\partial_t \chi)^2 dx.$$

So replacing in (2.15) we obtain (2.13).

We differentiate (2.9) and (2.11) with respect to t :

$$(i) \quad (1 + \chi) \partial_{tt} E + 2 \partial_t \chi \partial_t E + \partial_{tt} \chi E - \text{curl } \partial_t H = \partial_t G_1,$$

$$(ii) \quad \partial_{tt} H + \text{curl } \partial_t E = \partial_t G_2,$$

$$(iii) \quad \partial_{tt} \chi = \frac{2}{\varepsilon} E \cdot \partial_t E - \frac{1}{\varepsilon} \partial_t \chi,$$

$$\partial_t H \times n + a((\partial_t E \times n) \times n) = 0 \text{ in } \mathbb{R}^+ \times \Gamma.$$

As before we obtain

$$\begin{aligned} \frac{1}{2} \frac{d}{dt} \int_{\Omega} ((1 + \chi)|\partial_t E|^2 + |\partial_t H|^2) dx + \frac{3}{2} \int_{\Omega} \partial_t \chi |\partial_t E|^2 dx + \int_{\Omega} \partial_{tt} \chi E \cdot \partial_t E dx + \int_{\Gamma} a(\partial_t E_T) \cdot \partial_t E_T dx_2 dx_3 \\ = \int_{\Omega} (\partial_t G_1 \cdot \partial_t E + \partial_t G_2 \cdot \partial_t H) dx. \end{aligned}$$

The last equation in (2.9) yields

$$\int_{\Omega} \partial_{tt} \chi E \cdot \partial_t E dx = \frac{1}{4} \frac{d}{dt} \int_{\Omega} (\partial_t \chi)^2 dx + \frac{\varepsilon}{2} \int_{\Omega} (\partial_{tt} \chi)^2 dx,$$

so (2.14) holds.

Remark 2 *It is clear that (2.14) holds also for the Cauchy problem.*

2.4 Global existence for the 2d TE Kerr-Debye system.

For the two-dimensional TE model, the system is the following:

$$\begin{cases} \partial_t D_2 - \partial_3 H_1 + \partial_1 H_3 = 0, \\ \partial_t H_1 - \partial_3 E_2 = 0, \\ \partial_t H_3 + \partial_1 E_2 = 0, \\ \partial_t \chi = \frac{1}{\varepsilon} (|E_2|^2 - \chi), \end{cases} \quad (2.16)$$

for $t > 0$, for $(x_1, x_3) \in \mathbb{R}^+ \times \mathbb{R}$, together with zero initial data

$$D_2(0, x_1, x_3) = H_1(0, x_1, x_3) = H_3(0, x_1, x_3) = \chi(0, x_1, x_3) = 0 \text{ for } x_1 > 0, x_3 \in \mathbb{R}, \quad (2.17)$$

and the impedance boundary condition

$$H_3(t, 0, x_3) + aE_2(t, 0, x_3) = \varphi(t, x_3) \text{ for } t > 0, x_3 \in \mathbb{R}. \quad (2.18)$$

Recall that we have $\chi \geq 0$ and

$$\partial_1 H_1 + \partial_3 H_3 = 0. \quad (2.19)$$

We obtain the following global existence result.

Theorem 1 *Let $\varphi \in H^5(\mathbb{R}^2)$, compactly supported in $\mathbb{R}^+ \times \mathbb{R}$. Let $\varepsilon > 0$. Then the smooth solution $W = (D_2, H_1, H_3, \chi)$ to the IBVP (2.16)-(2.17)-(2.18) is defined on $[0, +\infty[$ ($T_{\varepsilon}^* = +\infty$).*

Proof of Theorem 1.

We fix $\varepsilon > 0$ and we assume that $T_{\varepsilon}^* < +\infty$. Then from Proposition 2, if $T_{\varepsilon}^* < +\infty$, then

$$\lim_{T \rightarrow T_{\varepsilon}^*} \|W\|_{\mathcal{H}^3(\Omega_T)} = +\infty, \quad (2.20)$$

where

$$\mathcal{H}^p(\Omega_T) = \left\{ W \text{ such that } \|W\|_{\mathcal{H}^p(\Omega_T)} := \sum_{i=0}^p \|\partial_t^i W\|_{L^\infty(0, T; H^{p-i}(\Omega))} < +\infty \right\}.$$

By variational estimates we will prove uniform bounds on $\|W\|_{\mathcal{H}^3(\Omega_T)}$ for $T < T_\varepsilon^*$ and so we will obtain a contradiction with (2.20), which proves that $T_\varepsilon^* = +\infty$.

With this object, as in the 3-d case, we extend the boundary condition (2.20): we replace $H_3(t, x_1, x_3)$ by $H_3(t, x_1, x_3) + \varphi(t, x_3)\eta(x_1)$ where η is a smooth function compactly supported in \mathbb{R}^+ , equal to 1 in a neighbourhood of 0. Furthermore, to preserve the divergence free condition (2.19) we replace $H_1(t, x_1, x_3)$

by $H_1(t, x_1, x_3) + \partial_3\varphi(t, x_3) \int_{x_1}^{+\infty} \eta(s)ds$. Then the system (2.16)-(2.17)-(2.18) becomes, in the variable $V = (E_2, H_1, H_3, \chi)$:

$$\begin{cases} (1 + \chi)\partial_t E_2 - \partial_3 H_1 + \partial_1 H_3 = -\partial_t \chi E_2 + G_1, \\ \partial_t H_1 - \partial_3 E_2 = G_2, \\ \partial_t H_3 + \partial_1 E_2 = G_3, \end{cases} \quad (2.21)$$

$$\partial_t \chi = \frac{1}{\varepsilon} (|E_2|^2 - \chi), \quad (2.22)$$

together with both zero initial data and homogeneous boundary condition:

$$V(0, x_1, x_3) = 0,$$

$$H_3(t, 0, x_3) + aE_2(t, 0, x_3) = 0, \quad (2.23)$$

where

$$G_1(t, x_1, x_3) = \partial_3^2 \varphi(t, x_3) \int_{x_1}^{+\infty} \eta(s)ds - \varphi(t, x_3)\eta'(x_1),$$

$$G_2(t, x_1, x_3) = \partial_t \partial_3 \varphi(t, x_3) \int_{x_1}^{+\infty} \eta(s)ds,$$

$$G_3(t, x_1, x_3) = -\partial_t \varphi(t, x_3)\eta(x_1).$$

We recall that the field H remains divergence free:

$$\partial_1 H_1 + \partial_3 H_3 = 0. \quad (2.24)$$

In this case the dissipation formulae (2.13)-(2.14) read

$$\frac{1}{2} \frac{d}{dt} \int_{\Omega} \left((1 + \chi)|E_2|^2 + |H|^2 + \frac{1}{2}|\chi|^2 \right) dx_1 dx_3 + \frac{\varepsilon}{2} \int_{\Omega} |\partial_t \chi|^2 dx + \int_{\Gamma} a|E_2|^2 dx_3 = \int_{\Omega} G \cdot U dx. \quad (2.25)$$

and

$$\begin{aligned} \frac{1}{2} \frac{d}{dt} \int_{\Omega} \left((1 + \chi)|\partial_t E_2|^2 + |\partial_t H|^2 + \frac{1}{2}|\partial_t \chi|^2 \right) dx_1 dx_3 + \frac{3}{2\varepsilon} \int_{\Omega} |E_2|^2 |\partial_t E_2|^2 dx + \frac{\varepsilon}{2} \int_{\Omega} |\partial_{tt} \chi|^2 dx \\ + \int_{\Gamma} a|\partial_t E_2|^2 dx_3 = \frac{3}{2\varepsilon} \int_{\Omega} \chi |\partial_t E|^2 dx + \int_{\Omega} \partial_t G \cdot \partial_t U dx \end{aligned} \quad (2.26)$$

where $G = (G_1, G_2, G_3)$ and $U = (E_2, H_1, H_3)$.

From (2.25), using Gronwall lemma (recall that $\chi \geq 0$), there exists a constant C such that

$$\|V\|_{\mathcal{H}^0(\Omega_{T_\varepsilon^*})} = \|V\|_{L^\infty(0, T_\varepsilon^*; L^2(\Omega))} \leq C. \quad (2.27)$$

Estimates on the first order time derivatives.

We recall the derivatives of (2.21)-(2.22)-(2.23) with respect to t :

$$\begin{cases} (1 + \chi)\partial_t^2 E_2 + 2\partial_t \chi \partial_t E_2 + \partial_t^2 \chi E_2 - \partial_3 \partial_t H_1 + \partial_1 \partial_t H_3 = \partial_t G_1, \\ \partial_t^2 H_1 - \partial_3 \partial_t E_2 = \partial_t G_2, \\ \partial_t^2 H_3 + \partial_1 \partial_t E_2 = \partial_t G_3, \end{cases} \quad (2.28)$$

$$\partial_t^2 \chi = \frac{1}{\varepsilon}(2E_2 \partial_t E_2 - \partial_t \chi), \quad (2.29)$$

$$\partial_t H_3(t, 0, x_3) + a \partial_t E_2(t, 0, x_3) = 0. \quad (2.30)$$

By the equations, we obtain zero initial data

$$\partial_t V(0, x_1, x_3) = 0.$$

From these equations, we obtained (2.26), and by Gronwall lemma, there exists $C > 0$ such that

$$\|\partial_t V\|_{L^\infty(0, T_\varepsilon^*; L^2(\Omega))} \leq C. \quad (2.31)$$

The last two equations in (2.21) yield

$$\|E_2\|_{L^\infty(0, T_\varepsilon^*; H^1(\Omega))} \leq C. \quad (2.32)$$

Solving (2.22) we get

$$\chi(t, x) = \frac{1}{\varepsilon} \int_0^t \exp\left(\frac{s-t}{\varepsilon}\right) |E_2(s, x)|^2 ds. \quad (2.33)$$

From Sobolev theorem and (2.32), for all $p \in [2, +\infty[$ there exists $C > 0$ such that

$$\|E_2\|_{L^\infty(0, T_\varepsilon^*; L^p(\Omega))} \leq C, \quad (2.34)$$

and so with (2.33), for all p with $1 \leq p < +\infty$, there exists C such that

$$\|\chi\|_{L^\infty(0, T_\varepsilon^*; L^p(\Omega))} \leq C. \quad (2.35)$$

So by (2.22) we obtain that, for all $p \in [1, +\infty[$,

$$\|\partial_t \chi\|_{L^\infty(0, T_\varepsilon^*; L^p(\Omega))} \leq C. \quad (2.36)$$

Estimates on the second order time derivatives.

We differentiate (2.28)-(2.29)-(2.30) with respect to t . We obtain

$$\begin{cases} (1 + \chi)\partial_t^3 E_2 + 3\partial_t \chi \partial_t^2 E_2 + 3\partial_t^2 \chi \partial_t E_2 + \partial_t^3 \chi E_2 - \partial_3 \partial_t^2 H_1 + \partial_1 \partial_t^2 H_3 = \partial_t^2 G_1, \\ \partial_t^3 H_1 - \partial_3 \partial_t^2 E_2 = \partial_t^2 G_2, \\ \partial_t^3 H_3 + \partial_1 \partial_t^2 E_2 = \partial_t^2 G_3, \end{cases} \quad (2.37)$$

$$\partial_t^3 \chi = \frac{1}{\varepsilon}(2|\partial_t E_2|^2 + 2E_2 \partial_t^2 E_2 - \partial_t^2 \chi), \quad (2.38)$$

with homogeneous boundary condition

$$\partial_t^2 H_3(t, 0, x_3) + a \partial_t^2 E_2(t, 0, x_3) = 0, \quad (2.39)$$

and by the equations, the initial data vanish

$$\partial_t^2 V(0, x_1, x_3) = 0.$$

Taking the inner product of (2.37) with $\partial_t^2 U$ we obtain

$$\begin{aligned} & \frac{1}{2} \frac{d}{dt} \int_{\Omega} ((1 + \chi)|\partial_t^2 E_2|^2 + |\partial_t^2 H|^2) dx_1 dx_3 + \int_{\Gamma} a |\partial_t^2 E_2|^2 dx_3 \\ & + \frac{5}{2} \int_{\Omega} \partial_t \chi |\partial_t^2 E_2|^2 dx + 3 \int_{\Omega} \partial_t^2 \chi \partial_t E_2 \partial_t^2 E_2 + \int_{\Omega} \partial_t^3 \chi E_2 \partial_t^2 E_2 = \int_{\Omega} \partial_t^2 G \cdot \partial_t^2 U dx. \end{aligned}$$

Equation (2.22) yields

$$\int_{\Omega} \partial_t \chi |\partial_t^2 E_2|^2 dx = \frac{1}{\varepsilon} \int_{\Omega} |E_2|^2 |\partial_t^2 E_2|^2 dx - \frac{1}{\varepsilon} \int_{\Omega} \chi |\partial_t^2 E_2|^2 dx.$$

From (2.29) we get

$$\int_{\Omega} \partial_t^2 \chi \partial_t E_2 \partial_t^2 E_2 dx = \frac{2}{\varepsilon} \int_{\Omega} E_2 |\partial_t E_2|^2 \partial_t^2 E_2 dx - \frac{1}{\varepsilon} \int_{\Omega} \partial_t \chi \partial_t E_2 \partial_t^2 E_2 dx.$$

From (2.38) we have

$$\begin{aligned} \int_{\Omega} \partial_t^3 \chi E_2 \partial_t^2 E_2 dx &= \frac{2}{\varepsilon} \int_{\Omega} E_2 |\partial_t E_2|^2 \partial_t^2 E_2 dx + \frac{2}{\varepsilon} \int_{\Omega} |E_2|^2 |\partial_t^2 E_2|^2 dx - \frac{2}{\varepsilon^2} \int_{\Omega} |E_2|^2 \partial_t E_2 \partial_t^2 E_2 dx \\ & \quad + \frac{1}{\varepsilon^2} \int_{\Omega} \partial_t \chi E_2 \partial_t^2 E_2 dx. \end{aligned}$$

So we obtain

$$\begin{aligned} & \frac{1}{2} \frac{d}{dt} \int_{\Omega} ((1 + \chi)|\partial_t^2 E_2|^2 + |\partial_t^2 H|^2) dx_1 dx_3 + \int_{\Gamma} a |\partial_t^2 E_2|^2 dx_3 + \frac{9}{2\varepsilon} \int_{\Omega} |E_2|^2 |\partial_t^2 E_2|^2 dx = \\ & \int_{\Omega} \partial_t^2 G \cdot \partial_t^2 U dx + \frac{5}{2\varepsilon} \int_{\Omega} \chi |\partial_t^2 E_2|^2 dx - \frac{8}{\varepsilon} \int_{\Omega} E_2 \partial_t^2 E_2 |\partial_t E_2|^2 dx + \frac{3}{\varepsilon} \int_{\Omega} \partial_t \chi \partial_t E_2 \partial_t^2 E_2 dx \\ & \quad + \frac{2}{\varepsilon^2} \int_{\Omega} |E_2|^2 \partial_t E_2 \partial_t^2 E_2 dx - \frac{1}{\varepsilon^2} \int_{\Omega} \partial_t \chi E_2 \partial_t^2 E_2 dx. \end{aligned} \quad (2.40)$$

By (2.28) and (2.31) we have

$$\|\partial_t E_2\|_{H^1(\Omega)} \leq C(1 + \|\partial_t^2 H\|_{L^2(\Omega)}). \quad (2.41)$$

Let us estimate the right hand side terms in (2.40). First we have

$$\left| \frac{8}{\varepsilon} \int_{\Omega} E_2 \partial_t^2 E_2 |\partial_t E_2|^2 dx \right| \leq \frac{9}{4\varepsilon} \int_{\Omega} |E_2|^2 |\partial_t^2 E_2|^2 dx + K(\varepsilon) \int_{\Omega} |\partial_t E_2|^4 dx.$$

By interpolation inequalities and Sobolev theorem, we have

$$\|u\|_{L^4(\Omega)} \leq C \|u\|_{H^{\frac{1}{2}}(\Omega)} \leq \|u\|_{L^2(\Omega)}^{\frac{1}{2}} \|u\|_{H^1(\Omega)}^{\frac{1}{2}}.$$

So using (2.41)

$$\left| \frac{8}{\varepsilon} \int_{\Omega} E_2 \partial_t^2 E_2 |\partial_t E_2|^2 dx \right| \leq \frac{9}{4\varepsilon} \int_{\Omega} |E_2|^2 |\partial_t^2 E_2|^2 dx + K(\varepsilon)(1 + \|\partial_t^2 H\|_{L^2(\Omega)}^2). \quad (2.42)$$

By (2.36) and (2.41) we obtain

$$\begin{aligned} \left| \int_{\Omega} \partial_t \chi \partial_t E_2 \partial_t^2 E_2 dx \right| &\leq C \|\partial_t \chi\|_{L^4(\Omega)} \|\partial_t E_2\|_{L^4(\Omega)} \|\partial_t^2 E_2\|_{L^2(\Omega)} \\ &\leq C \|\partial_t E_2\|_{H^1(\Omega)} \|\partial_t^2 E_2\|_{L^2(\Omega)} \\ &\leq C(1 + \|\partial_t^2 U\|_{L^2(\Omega)}^2). \end{aligned} \quad (2.43)$$

In the same way, (2.34) and (2.41) yield

$$\left| \int_{\Omega} |E_2|^2 \partial_t E_2 \partial_t^2 E_2 dx \right| \leq C(1 + \|\partial_t^2 U\|_{L^2(\Omega)}^2), \quad (2.44)$$

and by (2.34) and (2.36) we have

$$\left| \int_{\Omega} \partial_t \chi E_2 \partial_t^2 E_2 dx \right| \leq C(1 + \|\partial_t^2 U\|_{L^2(\Omega)}^2). \quad (2.45)$$

Using (2.42), (2.43), (2.44), (2.45) in (2.40), we obtain

$$\frac{1}{2} \frac{d}{dt} \int_{\Omega} ((1 + \chi)|\partial_t^2 E_2|^2 + |\partial_t^2 H|^2) dx_1 dx_3 \leq C(1 + \|\partial_t^2 U\|_{L^2(\Omega)}^2) + \frac{5}{2\varepsilon} \int_{\Omega} \chi |\partial_t^2 E_2|^2 dx,$$

so by Gronwall lemma there exists $C = C(T_\varepsilon^*)$ such that

$$\|\partial_t^2 U\|_{L^\infty(0, T_\varepsilon^*; L^2(\Omega))} \leq C. \quad (2.46)$$

So by (2.41)

$$\|\partial_t E_2\|_{L^\infty(0, T_\varepsilon^*; H^1(\Omega))} \leq C, \quad (2.47)$$

and by Sobolev theorem, for all p , $2 \leq p < +\infty$,

$$\|\partial_t E_2\|_{L^\infty(0, T_\varepsilon^*; L^p(\Omega))} \leq C. \quad (2.48)$$

As $\partial_t^2 \chi = \frac{2}{\varepsilon} E_2 \partial_t E_2 - \frac{1}{\varepsilon} \partial_t \chi$, by (2.34), (2.36), (2.48) we have for $1 \leq p < +\infty$:

$$\|\partial_t^2 \chi\|_{L^\infty(0, T_\varepsilon^*; L^p(\Omega))} \leq C. \quad (2.49)$$

Estimates on the third order time derivatives.

We differentiate (2.37)-(2.38)-(2.39) with respect to t :

$$\begin{cases} (1 + \chi) \partial_t^4 E_2 + 4 \partial_t \chi \partial_t^3 E_2 + 6 \partial_t^2 \chi \partial_t^2 E_2 + 4 \partial_t^3 \chi \partial_t E_2 + \partial_t^4 \chi E_2 - \partial_3 \partial_t^3 H_1 + \partial_1 \partial_t^3 H_3 = \partial_t^3 G_1, \\ \partial_t^4 H_1 - \partial_3 \partial_t^3 E_2 = \partial_t^3 G_2, \\ \partial_t^4 H_3 + \partial_1 \partial_t^3 E_2 = \partial_t^3 G_3, \end{cases} \quad (2.50)$$

$$\partial_t^4 \chi = \frac{6}{\varepsilon} \partial_t E_2 \partial_t^2 E_2 + \frac{2}{\varepsilon} E_2 \partial_t^3 E_2 - \frac{2}{\varepsilon^2} |\partial_t E_2|^2 - \frac{2}{\varepsilon^2} E_2 \partial_t^2 E_2 + \frac{1}{\varepsilon^2} \partial_t^2 \chi, \quad (2.51)$$

together with homogeneous boundary condition:

$$\partial_t^3 H_3(t, 0, x_3) + a \partial_t^3 E_2(t, 0, x_3) = 0,$$

and by the equations, the initial data vanish:

$$\partial_t^3 V(0, x_1, x_3) = 0.$$

We take the inner product of (2.50) with $\partial_t^3 U$ and we obtain

$$\begin{aligned} & \frac{1}{2} \frac{d}{dt} \int_{\Omega} ((1 + \chi)|\partial_t^3 E_2|^2 + |\partial_t^3 H|^2) dx_1 dx_3 + \int_{\Gamma} a |\partial_t^3 E_2|^2 dx_3 + \frac{7}{2} \int_{\Omega} \partial_t \chi |\partial_t^3 E_2|^2 dx + 6 \int_{\Omega} \partial_t^2 \chi \partial_t^2 E_2 \partial_t^3 E_2 \\ & + 4 \int_{\Omega} \partial_t^3 \chi \partial_t E_2 \partial_t^3 E_2 + \int_{\Omega} \partial_t^4 \chi E_2 \partial_t^3 E_2 = \int_{\Omega} \partial_t^3 G \cdot \partial_t^3 U dx. \end{aligned}$$

Equation (2.22) yields

$$\int_{\Omega} \partial_t \chi |\partial_t^3 E_2|^2 dx = \frac{1}{\varepsilon} \int_{\Omega} |E_2|^2 |\partial_t^3 E_2|^2 dx - \frac{1}{\varepsilon} \int_{\Omega} \chi |\partial_t^3 E_2|^2 dx.$$

From (2.38)

$$\int_{\Omega} \partial_t^3 \chi \partial_t E_2 \partial_t^3 E_2 = \frac{2}{\varepsilon} \int_{\Omega} |\partial_t E_2|^3 \partial_t^3 E_2 dx + \frac{2}{\varepsilon} \int_{\Omega} E_2 \partial_t E_2 \partial_t^2 E_2 \partial_t^3 E_2 dx - \int_{\Omega} \partial_t^2 \chi \partial_t E_2 \partial_t^3 E_2.$$

From (2.51), we get

$$\begin{aligned} \int_{\Omega} \partial_t^4 \chi E_2 \partial_t^3 E_2 &= \frac{6}{\varepsilon} \int_{\Omega} E_2 \partial_t E_2 \partial_t^2 E_2 \partial_t^3 E_2 dx + \frac{2}{\varepsilon} \int_{\Omega} |E_2|^2 |\partial_t^3 E_2|^2 dx - \frac{2}{\varepsilon^2} \int_{\Omega} E_2 |\partial_t E_2|^2 \partial_t^3 E_2 dx \\ &\quad - \frac{2}{\varepsilon^2} \int_{\Omega} |E_2|^2 \partial_t^2 E_2 \partial_t^3 E_2 dx + \frac{1}{\varepsilon^2} \int_{\Omega} \partial_t^2 \chi E_2 \partial_t^3 E_2. \end{aligned}$$

So we arrive at

$$\begin{aligned} \frac{1}{2} \frac{d}{dt} \int_{\Omega} ((1 + \chi) |\partial_t^3 E_2|^2 + |\partial_t^3 H|^2) dx_1 dx_3 &+ \int_{\Gamma} a |\partial_t^3 E_2|^2 dx_3 + \frac{11}{2\varepsilon} \int_{\Omega} |E_2|^2 |\partial_t^3 E_2|^2 dx \\ &= \int_{\Omega} \partial_t^3 G \cdot \partial_t^3 U dx + \frac{7}{2\varepsilon} \int_{\Omega} \chi |\partial_t^3 E_2|^2 dx + I_1 + I_2, \end{aligned} \quad (2.52)$$

where

$$\begin{aligned} I_1 &= -6 \int_{\Omega} \partial_t^2 \chi \partial_t^2 E_2 \partial_t^3 E_2 dx - \frac{14}{\varepsilon} \int_{\Omega} E_2 \partial_t E_2 \partial_t^2 E_2 \partial_t^3 E_2 dx + \frac{2}{\varepsilon} \int_{\Omega} |E_2|^2 \partial_t^2 E_2 \partial_t^3 E_2 dx, \\ I_2 &= -\frac{8}{\varepsilon} \int_{\Omega} |\partial_t E_2|^3 \partial_t^3 E_2 dx + \frac{4}{\varepsilon} \int_{\Omega} \partial_t^2 \chi \partial_t E_2 \partial_t^3 E_2 dx + \frac{2}{\varepsilon^2} \int_{\Omega} E_2 |\partial_t E_2|^2 \partial_t^3 E_2 dx \\ &\quad - \frac{1}{\varepsilon^2} \int_{\Omega} \partial_t^2 \chi E_2 \partial_t^3 E_2 dx. \end{aligned}$$

From (2.37) and (2.46) we have

$$\|\partial_t^2 E_2\|_{H^1(\Omega)} \leq K(1 + \|\partial_t^3 H\|_{L^2(\Omega)}). \quad (2.53)$$

So by Sobolev theorem, (2.34), (2.48), (2.49) and (2.53), we get

$$|I_1| \leq C(1 + \|\partial_t^3 U\|_{L^2(\Omega)}^2). \quad (2.54)$$

By (2.34), (2.48), (2.49), we have

$$|I_2| \leq C\|\partial_t^3 E_2\|_{L^2(\Omega)}. \quad (2.55)$$

Estimates (2.54) and (2.55) together with (2.52) imply

$$\frac{1}{2} \frac{d}{dt} \int_{\Omega} ((1 + \chi) |\partial_t^3 E_2|^2 + |\partial_t^3 H|^2) dx_1 dx_3 \leq C(1 + \|\partial_t^3 U\|_{L^2(\Omega)}^2) + \frac{7}{2\varepsilon} \int_{\Omega} \chi |\partial_t^3 E_2|^2 dx.$$

So there exists a constant $C = C(T_\varepsilon^*) > 0$ such that

$$\|\partial_t^3 U\|_{L^\infty(0, T_\varepsilon^*; L^2(\Omega))} \leq C.$$

By (2.53)

$$\|\partial_t^2 E_2\|_{L^\infty(0, T_\varepsilon^*; H^1(\Omega))} \leq C,$$

and, for all $p, 2 \leq p < +\infty$,

$$\|\partial_t^2 E_2\|_{L^\infty(0, T_\varepsilon^*; L^p(\Omega))} \leq C. \quad (2.56)$$

As $\partial_t^3 \chi = \frac{2}{\varepsilon} |\partial_t E_2|^2 + \frac{2}{\varepsilon} E_2 \partial_t^2 E_2 - \frac{1}{\varepsilon} \partial_t^2 \chi$, using (2.34), (2.48), (2.49) and (2.56), for $1 \leq p < +\infty$,

$$\|\partial_t^3 \chi\|_{L^\infty(0, T_\varepsilon^*; L^p(\Omega))} \leq C.$$

Estimates in $\mathcal{H}^3(\Omega_{T_\varepsilon^*})$.

In order to estimate the space derivatives of H we recall the following div-curl lemma (see [9]).

Lemma 1 *Let $\Omega = \{(x_1, x_3), x_1 > 0\}$. We denote $\Gamma = \partial\Omega$. Let $H = (H_1, H_3) \in H^k(\Omega)$ such that $\operatorname{div} H = \partial_1 H_1 + \partial_3 H_3 \in H^k(\Omega)$, $\operatorname{curl} H = -\partial_1 H_3 + \partial_3 H_1 \in H^k(\Omega)$, $H_3 \in H^{k+\frac{1}{2}}(\Gamma)$. Then $H \in H^{k+1}(\Omega)$ and we have*

$$\|H\|_{H^{k+1}(\Omega)} \leq C \left(\|H\|_{H^k(\Omega)} + \|\operatorname{div} H\|_{H^k(\Omega)} + \|\operatorname{curl} H\|_{H^k(\Omega)} + \|H\|_{H^{k+\frac{1}{2}}(\Gamma)} \right).$$

By the first equation in (2.21), using (2.34), (2.35), (2.36) and (2.48), we obtain

$$\|\operatorname{curl} H\|_{L^\infty(0, T_\varepsilon^*; L^2(\Omega))} \leq C.$$

Recall that $\operatorname{div} H = 0$ and on the boundary Γ , $H_3(t, 0, x_3) = -aE_2(t, 0, x_3)$, so by (2.32),

$$\|H_3(t, 0, \cdot)\|_{L^\infty(0, T_\varepsilon^*; H^{\frac{1}{2}}(\Gamma))} \leq C.$$

Using lemma 1, we obtain

$$\|H\|_{L^\infty(0, T_\varepsilon^*; H^1(\Omega))} \leq C.$$

By the first equation in (2.28), using (2.34), (2.35), (2.36), (2.48), (2.49) and (2.56), we get

$$\|\operatorname{curl} \partial_t H\|_{L^\infty(0, T_\varepsilon^*; L^2(\Omega))} \leq C.$$

Condition (2.23) together with (2.47) yield

$$\|\partial_t H_3(t, 0, \cdot)\|_{L^\infty(0, T_\varepsilon^*; H^{\frac{1}{2}}(\Gamma))} \leq C,$$

thus

$$\|\partial_t H\|_{L^\infty(0, T_\varepsilon^*; H^1(\Omega))} \leq C. \quad (2.57)$$

By the second and the third equations in (2.21), by (2.32) and (2.57),

$$\|E_2\|_{L^\infty(0, T_\varepsilon^*; H^2(\Omega))} \leq C, \quad (2.58)$$

and by Sobolev theorem,

$$\|E_2\|_{L^\infty([0, T_\varepsilon^*] \times \Omega)} \leq C.$$

As $H^2(\Omega)$ is an algebra, using (2.33), we obtain

$$\|\chi\|_{L^\infty(0, T_\varepsilon^*; H^2(\Omega))} \leq C, \quad (2.59)$$

and by Sobolev theorem,

$$\|\chi\|_{L^\infty([0, T_\varepsilon^*] \times \Omega)} \leq C.$$

Since $\partial_t \chi = \frac{1}{\varepsilon} (|E_2|^2 - \chi)$,

$$\|\partial_t \chi\|_{L^\infty(0, T_\varepsilon^*; H^2(\Omega))} \leq C, \quad (2.60)$$

and by Sobolev theorem,

$$\|\partial_t \chi\|_{L^\infty([0, T_\varepsilon^*] \times \Omega)} \leq C.$$

In addition with (2.38) we have

$$\|\partial_t^2 \chi\|_{L^\infty(0, T_\varepsilon^*; H^1(\Omega))} \leq C.$$

By the first equation in (2.21), using (2.47), (2.58), (2.59), (2.60), we obtain

$$\|\operatorname{curl} H\|_{L^\infty(0, T_\varepsilon^*; H^1(\Omega))} \leq C.$$

Using (2.58), (2.23) yields

$$\|H_3(t, 0, \cdot)\|_{L^\infty(0, T_\varepsilon^*; H^{\frac{3}{2}}(\Gamma))} \leq C,$$

so by Lemma 1,

$$\|H\|_{L^\infty(0, T_\varepsilon^*; H^2(\Omega))} \leq C.$$

By analogous arguments, we prove successively that

$$\|\partial_t^2 H\|_{L^\infty(0, T_\varepsilon^*; H^1(\Omega))} \leq C,$$

$$\|\partial_t E_2\|_{L^\infty(0, T_\varepsilon^*; H^2(\Omega))} \leq C,$$

$$\|\partial_t^2 \chi\|_{L^\infty(0, T_\varepsilon^*; H^2(\Omega))} \leq C,$$

$$\|\partial_t H\|_{L^\infty(0, T_\varepsilon^*; H^2(\Omega))} \leq C,$$

$$\|E_2\|_{L^\infty(0, T_\varepsilon^*; H^3(\Omega))} \leq C,$$

$$\|\chi\|_{L^\infty(0, T_\varepsilon^*; H^3(\Omega))} \leq C,$$

$$\|H\|_{L^\infty(0, T_\varepsilon^*; H^3(\Omega))} \leq C.$$

Finally,

$$\|V\|_{\mathcal{H}^3(\Omega_{T_\varepsilon^*})} \leq C.$$

This estimate contradicts (2.20), so $T_\varepsilon^* = +\infty$. The proof of Theorem 1 is complete.

2.5 Global existence for the 1-d Kerr-Debye system.

The one-dimensional Kerr Debye IBVP writes:

$$\begin{cases} \partial_t d + \partial_x h = 0, \\ \partial_t h + \partial_x e = 0, \\ \partial_t \chi = \frac{1}{\varepsilon}(e^2 - \chi), \\ d = (1 + \chi)e, \end{cases} \quad (2.61)$$

for $(t, x) \in \mathbb{R}^+ \times \mathbb{R}^+$, together with both zero initial data

$$(d, h, \chi)(0, x) = 0 \text{ for } x \geq 0, \quad (2.62)$$

and the boundary condition

$$h(t, 0) + ae(t, 0) = \varphi(t) \text{ for } t \geq 0. \quad (2.63)$$

Theorem 2 *Let $\varphi \in H^4(\mathbb{R})$ compactly supported in \mathbb{R}^+ . Let $\varepsilon > 0$. Then the smooth solution (d, h, χ) to the IBVP (2.61)-(2.62)-(2.63) is defined on $[0, +\infty[$.*

The proof of Theorem 2 is analogous to the proof of Theorem 1. The one-dimensional case is easier: the divergence free condition is irrelevant and the space regularity is obtained directly by the equations. In addition, by Sobolev theorem, the H^1 -estimates implies bounds in the L^∞ -norm. We refer to [7] for the detailed proof of the global existence for the Cauchy problem.

Remark 3 *Using the general stability results in [12] and [1], we can observe that for the 1-d Kerr-Debye model, a constant equilibrium state $(\bar{e}, \bar{h}, \bar{\chi} = (\bar{e})^2)$ is stable if $\bar{e} \neq 0$. The stability property if $\bar{e} = 0$ in the one-dimensional case and in the two-dimensional TE case remains an open problem.*

Remark 4 *It seems more difficult to obtain the global existence for both the 2-dTM and the 3-d models. First we lack for an adapted div-curl lemma to obtain the space regularity. Indeed, in both cases, the divergence free condition (2.7) is nonlinear in (χ, E) . In addition, the Sobolev embeddings used in the 2-d case don't work in the 3-d case.*

3 Convergence result.

First we replace the Kerr-Debye model in the general framework of [8]. The equilibrium manifold in (2.4) is defined by:

$$\mathcal{V} = \{(D, H, \chi) \in \mathbb{R}^7, \chi = |E|^2 = (1 + \chi)^{-2}|D|^2\},$$

so that the reduced system associated with the Kerr-Debye model (2.4) is the Kerr model (2.1).

The strictly convex entropy \mathcal{E}_{KD} satisfies the stability condition in the definition 2.1 of [8] and, on the equilibrium manifold \mathcal{V} , we have the relation:

$$\mathcal{E}_K(D, H) = \mathcal{E}_{KD}(D, H, \chi(D)).$$

Furthermore characteristic speeds associated to (2.4) and (2.1) are interlaced on the equilibrium manifold \mathcal{V} . In our case, by the previous calculations, in the 3-d model we have

$$\mu_1(|E|^2, \xi) = \lambda_1(E, \xi) = \mu_2(|E|^2, \xi) \leq \lambda_2(E, \xi) \leq 0 = \mu_3 = \lambda_3 = \mu_4 = \lambda_4 = \mu_5.$$

So the Kerr-Debye model is a relaxation approximation of the Kerr model, with the stability properties. In this context, it is natural to study the convergence of the solutions to the Kerr-Debye model as the relaxation coefficient tends to zero. The Cauchy problem is studied in [11]. Here we establish the convergence for the impedance IBVP.

3.1 Entropic variables and main result.

In order to prove the convergence result it is more convenient to use the entropic variables which are introduced in [12]. These variables are obtained taking the gradient of the convex entropy (2.8):

$$\begin{cases} \partial_D \mathcal{E}_{KD} = (1 + \chi)^{-1} D = E, \\ \partial_H \mathcal{E}_{KD} = H, \\ \partial_\chi \mathcal{E}_{KD} = \frac{1}{2}(\chi - |E|^2) := v, \end{cases}$$

The IBVP (2.4)-(2.5)-(2.6) becomes (for $(t, x) \in \mathbb{R}^+ \times \Omega$):

$$A_0(W_\varepsilon) \partial_t W_\varepsilon + \sum_{j=1}^3 A_j \partial_j W_\varepsilon = \frac{1}{\varepsilon} Q(W_\varepsilon), \quad (3.1)$$

where

- $W_\varepsilon = \begin{pmatrix} E_\varepsilon \\ H_\varepsilon \\ v_\varepsilon \end{pmatrix}, \quad Q(W_\varepsilon) = \begin{pmatrix} 0 \\ 0 \\ -2v_\varepsilon \end{pmatrix},$
- $A_0(W_\varepsilon) = \begin{pmatrix} (|E_\varepsilon|^2 + 2v_\varepsilon + 1)I_3 + 2E_\varepsilon {}^t E_\varepsilon & 0 & 2E_\varepsilon \\ 0 & I_3 & 0 \\ 2 {}^t E_\varepsilon & 0 & 2 \end{pmatrix},$
- $\sum_{j=1}^3 A_j \partial_j = \begin{pmatrix} 0 & -\text{curl} & 0 \\ \text{curl} & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix},$

with the initial data

$$E_\varepsilon(0, x) = H_\varepsilon(0, x) = 0, \quad v_\varepsilon(0, x) = 0 \text{ for } x \in \Omega, \quad (3.2)$$

and with the boundary condition

$$H_\varepsilon \times n + a((E_\varepsilon \times n) \times n) = \varphi \text{ for } (t, x) \in \mathbb{R}^+ \times \Gamma. \quad (3.3)$$

We observe that in the entropic variables (E, H, v) , the boundary condition is linear. In addition, with these variables, the equilibrium manifold is flat since $\mathcal{V} = \{(E, H, v) \in \mathbb{R}^7, v = 0\}$ and the relaxation term is linear.

We rewrite the divergence free condition in the entropic variables:

$$\text{div}((1 + |E_\varepsilon|^2 + 2v_\varepsilon)E_\varepsilon) = \text{div} H_\varepsilon = 0. \quad (3.4)$$

In the same way we can write the Kerr model in its entropic variables :

$$\begin{cases} \partial_D \mathcal{E}_K = E, \\ \partial_H \mathcal{E}_K = H. \end{cases}$$

The IBVP (2.1)-(2.2)-(2.3) becomes

$$\begin{cases} ((1 + |E_0|^2)I_3 + 2E_0 {}^t E_0)\partial_t E_0 - \text{curl} H_0 = 0, \\ \partial_t H_0 + \text{curl} H_0 = 0, \end{cases} \quad (3.5)$$

for $(t, x) \in \mathbb{R}^+ \times \Omega$ together with zero initial data

$$E_0(0, x) = H_0(0, x) = 0 \text{ for } x \in \Omega, \quad (3.6)$$

and with the impedance boundary condition

$$H_0 \times n + a((E_0 \times n) \times n) = \varphi \text{ for } (t, x) \in \mathbb{R}^+ \times \Gamma. \quad (3.7)$$

We remark that the divergence free condition is satisfied

$$\text{div}((1 + |E_0|^2)E_0) = \text{div} H_0 = 0. \quad (3.8)$$

In these entropic variables the IBVP (3.5)-(3.6)-(3.7) is the reduced system of the IBVP (3.1)-(3.2)-(3.3).

Let us recall the assumptions on the source term φ in both problems:

$$\begin{aligned} &\varphi \text{ is compactly supported in } \mathbb{R}^+ \times \Gamma, \\ &\varphi \in H^s(\mathbb{R} \times \Gamma), \quad s \geq 5. \end{aligned} \quad (3.9)$$

Thus we can apply Proposition 1 and Proposition 2 with the source term φ to define the solution (E_0, H_0) to (3.5)-(3.6)-(3.7) on $[0, T^*[$ and the solution $(E_\varepsilon, H_\varepsilon, v_\varepsilon)$ to (3.1)-(3.2)-(3.3) on $[0, T_\varepsilon^*[$. In addition we choose s great enough to ensure a sufficient regularity on the profile (E_0, H_0) (see Remark 1).

As in the previous section we introduce for $T > 0$ and $k \in \mathbb{N}$:

$$\mathcal{H}^k(\Omega_T) = \left\{ W \text{ such that } \|W\|_{\mathcal{H}^k(\Omega_T)} := \sum_{i=0}^k \|\partial_t^i W\|_{L^\infty(0, T; H^{k-i}(\Omega))} < +\infty \right\}.$$

We obtain the following convergence result.

Theorem 3 *Let φ satisfying (3.9). Let (E_0, H_0) be the solution to (3.5)-(3.6)-(3.7) which lifespan is denoted by T^* . We assume that for all $T < T^*$, $E_0 \in \mathcal{H}^7(\Omega_T)$. For $\varepsilon > 0$, let $(E_\varepsilon, H_\varepsilon, v_\varepsilon)$ be the solution to (3.1)-(3.2)-(3.3) which lifespan is denoted by T_ε^* . We fix $T < T^*$. There exists $\varepsilon_0 > 0$ such that*

$$\forall \varepsilon < \varepsilon_0, T_\varepsilon^* \geq T, \quad (3.10)$$

and there exists a constant C such that for all $\varepsilon < \varepsilon_0$,

$$\|E_\varepsilon - E_0\|_{\mathcal{H}^3(\Omega_T)} + \|H_\varepsilon - H_0\|_{\mathcal{H}^3(\Omega_T)} + \|v_\varepsilon\|_{\mathcal{H}^3(\Omega_T)} \leq C\varepsilon. \quad (3.11)$$

As remarked in the introduction, we don't expect boundary layer formation in the IBVP (3.1)-(3.2)-(3.3) near the profile $(E_0, H_0, 0)$. Hence we use a Hilbert expansion to describe the behavior of the solution as ε tends to zero.

3.2 Hilbert expansion.

We denote by $\rho_\varepsilon = {}^t(R_\varepsilon, S_\varepsilon, s_\varepsilon)$ the remainder term in the Hilbert expansion of W_ε :

$$E_\varepsilon = E_0 + \varepsilon R_\varepsilon,$$

$$H_\varepsilon = H_0 + \varepsilon S_\varepsilon,$$

$$v_\varepsilon = \varepsilon s_1 + \varepsilon s_\varepsilon,$$

with $s_1 = -E_0 \partial_t E_0$. Using (3.5)-(3.6)-(3.7) and (3.1)-(3.2)-(3.3), the rest term ρ_ε satisfies the following system:

$$A_0(t, x) \partial_t \rho_\varepsilon + \sum_{j=1}^3 A_j \partial_j \rho_\varepsilon + L(t, x) \rho_\varepsilon + B(t, x) + G(t, x, R_\varepsilon) + F(t, x, \tilde{\rho}_\varepsilon) \partial_t \tilde{\rho}_\varepsilon = -{}^t(0, 0, \frac{2}{\varepsilon} s_\varepsilon), \quad (3.12)$$

where

- $\tilde{\rho}_\varepsilon = \begin{pmatrix} R_\varepsilon \\ 0 \\ s_\varepsilon \end{pmatrix},$
- $A_0(t, x) = \begin{pmatrix} (1 + |E_0|^2)I_3 + 2E_0 {}^t E_0 & 0 & 2E_0 \\ 0 & I_3 & 0 \\ 2 {}^t E_0 & 0 & 2 \end{pmatrix},$
- $L(t, x) \rho_\varepsilon = \begin{pmatrix} 2s_\varepsilon \partial_t E_0 + 2E_0 \cdot R_\varepsilon \partial_t E_0 + 2(E_0 {}^t R_\varepsilon + R_\varepsilon {}^t E_0) \partial_t E_0 + 2\varepsilon \partial_t s_1 R_\varepsilon \\ 0 \\ 2 {}^t R_\varepsilon \partial_t E_0 \end{pmatrix},$
- $B(t, x) = \begin{pmatrix} 2s_1 \partial_t E_0 + 2\partial_t s_1 E_0 \\ 0 \\ 2\partial_t s_1 \end{pmatrix},$

- $G(t, x, R_\varepsilon) = \begin{pmatrix} \varepsilon |R_\varepsilon|^2 \partial_t E_0 + \varepsilon R_\varepsilon {}^t R_\varepsilon \partial_t E_0 \\ 0 \\ 0 \end{pmatrix},$
- $F(t, x, \tilde{\rho}_\varepsilon) \partial_t \tilde{\rho}_\varepsilon = \begin{pmatrix} \varepsilon (R_\varepsilon {}^t E_0 + E_0 {}^t R_\varepsilon) \partial_t R_\varepsilon + \varepsilon s_\varepsilon \partial_t R_\varepsilon + \varepsilon^2 R_\varepsilon {}^t R_\varepsilon \partial_t R_\varepsilon + 2\varepsilon R_\varepsilon \partial_t s_\varepsilon + \varepsilon^2 |R_\varepsilon|^2 \partial_t R_\varepsilon \\ 0 \\ 2\varepsilon R_\varepsilon \cdot \partial_t R_\varepsilon \end{pmatrix},$

with the zero initial data

$$\rho_\varepsilon(0, x) = 0 \text{ for } x \in \Omega, \quad (3.13)$$

and the homogeneous boundary condition

$$S_\varepsilon \times n + a((R_\varepsilon \times n) \times n) = 0 \text{ for } (t, x) \in \mathbb{R}^+ \times \Gamma. \quad (3.14)$$

We remark that we introduced the one order term εs_1 in the expansion of v_ε to avoid a singular source term in the last equation of System (3.12).

For $t \leq T_\varepsilon^*$ we define φ_ε and ψ_ε by

$$\varphi_\varepsilon(t) = \left(\|\rho_\varepsilon(t)\|_{L^2(\Omega)}^2 + \sum_{i \neq 1} \|\partial_i \rho_\varepsilon(t)\|_{L^2(\Omega)}^2 + \sum_{i, j \neq 1} \|\partial_{ij} \rho_\varepsilon(t)\|_{L^2(\Omega)}^2 + \sum_{i, j, k \neq 1} \|\partial_{ijk} \rho_\varepsilon(t)\|_{L^2(\Omega)}^2 \right)^{\frac{1}{2}},$$

$$\psi_\varepsilon(t) = \left(\|\partial_1 \rho_\varepsilon(t)\|_{L^2(\Omega)}^2 + \sum_i \|\partial_{1i} \rho_\varepsilon(t)\|_{L^2(\Omega)}^2 + \sum_{i, j} \|\partial_{1ij} \rho_\varepsilon(t)\|_{L^2(\Omega)}^2 \right)^{\frac{1}{2}}.$$

Let us remark that φ_ε measures the tangential derivatives $\partial_0 = \partial_t, \partial_2, \partial_3$ and we have

$$(\varphi_\varepsilon(t))^2 + (\psi_\varepsilon(t))^2 = \sum_{i=0}^3 \|\partial_t^i \rho_\varepsilon(t)\|_{H^{3-i}(\Omega)}^2.$$

We define also

$$\Phi_\varepsilon(t) = \sup_{s \in [0, t]} \varphi_\varepsilon(s),$$

$$\Psi_\varepsilon(t) = \sup_{s \in [0, t]} \psi_\varepsilon(s),$$

so we have

$$\Phi_\varepsilon(T) + \Psi_\varepsilon(T) \sim \|\rho_\varepsilon\|_{\mathcal{H}^3(\Omega_T)}.$$

3.3 Proof of Theorem 3.

We fix $T < T^*$ and we define T_ε by

$$T_\varepsilon = \sup \left\{ t \leq T, \|\rho_\varepsilon\|_{\mathcal{H}^3(\Omega_t)} \leq \frac{1}{\sqrt{\varepsilon}} \right\}, \quad (3.15)$$

so by Proposition 2, $T_\varepsilon \leq T_\varepsilon^*$.

The proof of Theorem 3 is organized as follows. In the first step, by variational methods on the system (3.12)-(3.13)-(3.14) we estimate the tangential derivatives of ρ_ε . In the second step we bound the normal derivatives of ρ_ε by solving the last equation in (3.12) and using the divergence free conditions (3.4) and (3.8).

We will use the following classical lemma.

Lemma 2 *For $k \geq 2$, $\mathcal{H}^k(\Omega_T)$ is an algebra.*

Since $E_0 \in \mathcal{H}^7(\Omega_T)$, we obtain

$$A_0, L, B \in \mathcal{H}^5(\Omega_T), \quad (3.16)$$

so A_0, L and B are sufficiently regular to allow the following calculations.

For the convenience of the reader, we will omit the dependence on the index ε .

First step: estimates on φ_ε .

Let us establish the following lemma:

Lemma 3 *There exists a constant C such that for $t \in [0, T_\varepsilon]$,*

$$\begin{aligned} \frac{d}{dt} \left[\int_{\Omega} \left(A_0 \rho \cdot \rho + \sum_{i \neq 1} A_0 \partial_i \rho \cdot \partial_i \rho + \sum_{i, j \neq 1} A_0 \partial_{ij} \rho \cdot \partial_{ij} \rho + \sum_{i, j, k \neq 1} A_0 \partial_{ijk} \rho \cdot \partial_{ijk} \rho \right) dx \right. \\ \left. + \int_{\Omega} \left(\sum_{i, j, k \neq 1} F(t, x, \tilde{\rho}) \partial_{ijk} \tilde{\rho} \cdot \partial_{ijk} \tilde{\rho} \right) dx \right] \leq C(1 + \varphi^2). \end{aligned} \quad (3.17)$$

Proof. The nonlinear terms in (3.12) are bounded thanks to the following estimates, which are a straight-forward consequence of Lemma 2 and (3.15): there exists a constant C such that for all $\varepsilon > 0$,

$$\begin{aligned} \|G(t, x, R)\|_{\mathcal{H}^3(\Omega_{T_\varepsilon})} &\leq C, \\ \|F(t, x, \tilde{\rho}) \partial_t \tilde{\rho}\|_{\mathcal{H}^2(\Omega_{T_\varepsilon})} &\leq C, \\ \|F(t, x, \tilde{\rho})\|_{\mathcal{H}^3(\Omega_{T_\varepsilon})} &\leq C\sqrt{\varepsilon}. \end{aligned} \quad (3.18)$$

Taking the inner product of (3.12) with ρ we obtain

$$\begin{aligned} \frac{1}{2} \frac{d}{dt} \int_{\Omega} A_0 \rho \cdot \rho dx - \int_{\Gamma} a((R \times n) \times n) \cdot R dx_2 dx_3 + \frac{2}{\varepsilon} \int_{\Omega} |s|^2 dx = \\ \frac{1}{2} \int_{\Omega} \partial_t A_0 \rho \cdot \rho dx - \int_{\Omega} L \rho \cdot \rho dx - \int_{\Omega} B \cdot \rho dx - \int_{\Omega} G(t, x, R) \cdot \rho dx - \int_{\Omega} F(t, x, \tilde{\rho}) \partial_t \tilde{\rho} \cdot \rho dx. \end{aligned}$$

By assumption on the operator a , we have

$$\int_{\Gamma} a((R \times n) \times n) \cdot R dx_2 dx_3 \leq 0.$$

Since E_0 is regular enough, by (3.16), we have immediately

$$\left| \frac{1}{2} \int_{\Omega} \partial_t A_0 \rho \cdot \rho dx - \int_{\Omega} L \rho \cdot \rho dx \right| \leq C(\varphi(t))^2,$$

and

$$\left| \int_{\Omega} B \cdot \rho dx \right| \leq C\varphi(t).$$

From (3.18),

$$\left| \int_{\Omega} G(t, x, R) \cdot \rho dx \right| \leq \|G\|_{L^2(\Omega)} \|\rho\|_{L^2(\Omega)} \leq \|G\|_{\mathcal{H}^3(\Omega_{T_\varepsilon})} \varphi(t) \leq C\varphi(t),$$

and

$$\left| \int_{\Omega} F(t, x, \tilde{\rho}) \partial_t \tilde{\rho} \cdot \rho dx \right| \leq \|F \partial_t \tilde{\rho}\|_{L^2(\Omega)} \|\rho\|_{L^2(\Omega)} \leq \|F \partial_t \tilde{\rho}\|_{\mathcal{H}^2(\Omega_{T_\varepsilon})} \|\rho\|_{L^2(\Omega)} \leq C\varphi(t).$$

Thus we obtain the L^2 estimate:

$$\frac{d}{dt} \int_{\Omega} A_0 \rho \cdot \rho dx \leq C(1 + (\varphi(t))^2).$$

In the same way, differentiating (3.12) with respect to the tangential variables, we obtain

$$\frac{d}{dt} \left[\int_{\Omega} \sum_{i \neq 1} A_0 \partial_i \rho \cdot \partial_i \rho dx + \int_{\Omega} \sum_{i, j \neq 1} A_0 \partial_{ij} \rho \cdot \partial_{ij} \rho dx \right] \leq C(1 + (\varphi(t))^2).$$

For the third order derivatives, we use also the estimate:
for $i, j, k \neq 1$,

$$\left| \int_{\Omega} \partial_{ijk} (F(t, x, \tilde{\rho}) \partial_t \tilde{\rho}) \cdot \partial_{ijk} \tilde{\rho} dx - \frac{1}{2} \frac{d}{dt} \int_{\Omega} F(t, x, \tilde{\rho}) \partial_{ijk} \tilde{\rho} \cdot \partial_{ijk} \tilde{\rho} dx \right| \leq C(1 + (\varphi(t))^2),$$

and we conclude the proof of Lemma 3.

Let us recall that $A_0(t, x) = A_0(E_0)$ (see (3.12)), so, since E_0 is smooth on $[0, T] \times \overline{\Omega}$, there exists a constant $\alpha > 0$ such that for all $\xi \in \mathbb{R}^3$ and $(t, x) \in [0, T] \times \overline{\Omega}$,

$$A_0(t, x) \xi \cdot \xi \geq \alpha |\xi|^2. \quad (3.19)$$

In addition, from the last estimate in (3.18) we remark that, for all $t \in [0, T_\varepsilon]$,

$$\left| \int_{\Omega} \sum_{i, j, k \neq 1} F(t, x, \tilde{\rho}) \partial_{ijk} \tilde{\rho} \cdot \partial_{ijk} \tilde{\rho} dx \right| \leq C_1 \sqrt{\varepsilon} (\varphi(t))^2. \quad (3.20)$$

So, integrating (3.17) and using (3.19) and (3.20) we obtain that, for $t \in [0, T_\varepsilon]$,

$$\alpha \varphi^2(t) - C_1 \sqrt{\varepsilon} \varphi^2(t) \leq C_2 \int_0^t (1 + \varphi^2(s)) ds,$$

so for ε small enough,

$$\varphi^2(t) \leq C \int_0^t (1 + \varphi^2(s)) ds,$$

so by Gronwall lemma, we obtain that there exists C such that for all $t \in [0, T_\varepsilon]$

$$\varphi^2(t) \leq C. \quad (3.21)$$

Second step: estimates on ψ_ε .

In order to estimate the derivatives with respect to the normal variable ∂_1 we proceed as follows:

1. We rewrite (3.12) isolating $\partial_1 R_3$, $\partial_1 R_2$, $\partial_1 S_3$ and $\partial_1 S_2$ in curl R and curl S :

$$\begin{cases} \partial_1 R_3 = \partial_3 R_1 + \partial_t S_2 \\ \partial_1 R_2 = \partial_2 R_1 - \partial_t S_3 \end{cases} \quad (3.22)$$

so from (3.21),

$$\|\partial_1 R_3\|_{L^\infty(0, T_\varepsilon; L^2(\Omega))} + \|\partial_1 R_2\|_{L^\infty(0, T_\varepsilon; L^2(\Omega))} \leq C, \quad (3.23)$$

and

$$\begin{cases} \partial_1 S_3 = \partial_3 S_1 - (A_0 \partial_t \rho)_2 - (L\rho)_2 + M_2(t, x, \varepsilon, \tilde{\rho}, \partial_t \tilde{\rho}), \\ \partial_1 S_2 = \partial_2 S_1 + (A_0 \partial_t \rho)_3 + (L\rho)_3 + M_3(t, x, \varepsilon, \tilde{\rho}, \partial_t \tilde{\rho}), \end{cases} \quad (3.24)$$

where, by (3.18),

$$\|M_2\|_{\mathcal{H}^2(\Omega_{T_\varepsilon})} + \|M_3\|_{\mathcal{H}^2(\Omega_{T_\varepsilon})} \leq C. \quad (3.25)$$

Using also (3.21), we obtain that

$$\|\partial_1 S_3\|_{L^\infty(0, T_\varepsilon; L^2(\Omega))} + \|\partial_1 S_2\|_{L^\infty(0, T_\varepsilon; L^2(\Omega))} \leq C.$$

2. In order to estimate $\partial_1 S_1$, $\partial_1 R_1$ and $\partial_1 s$, let us rewrite the divergence free conditions. From (3.4) and (3.8), we obtain

$$\partial_1 S_1 = -\partial_2 S_2 - \partial_3 S_3, \quad (3.26)$$

so we have

$$\|\partial_1 S_1\|_{L^\infty(0, T_\varepsilon; L^2(\Omega))} \leq C.$$

It remains to estimate $\partial_1 R_1$ and $\partial_1 s$. From (3.4) and (3.8), we have

$$\begin{aligned} \operatorname{div} \left((1 + |E_0|^2 R + 2(E_0 \cdot R)E_0 + 2sE_0) \right) &= \tilde{N}(t, x, \varepsilon \tilde{\rho}, \partial_1 \tilde{\rho}, \partial_2 \tilde{\rho}, \partial_3 \tilde{\rho}) \\ &:= -\operatorname{div} \left(2s_1 E_0 + \varepsilon |R|^2 E_0 + 2\varepsilon (s_1 + s + (E_0 \cdot R)R) + \varepsilon^2 |R|^2 R \right). \end{aligned}$$

We expand the left hand side term and we obtain

$$\begin{aligned} (1 + |E_0|^2 + 2(E_{0,1})^2) \partial_1 R_1 + 2E_{0,1} \partial_1 s + \tilde{\Lambda}(t, x) \rho + \tilde{\Lambda}_2(t, x) \partial_2 \rho + \tilde{\Lambda}_3(t, x) \partial_3 \rho \\ + 2E_{0,1} E_{0,2} \partial_1 R_2 + 2E_{0,1} E_{0,3} \partial_1 R_3 = \tilde{N} \end{aligned}$$

where $\tilde{\Lambda}(t, x)$, $\tilde{\Lambda}_2(t, x)$ and $\tilde{\Lambda}_3(t, x)$ are linear operators.

Using (3.22) we obtain

$$\partial_1 R_1 + 2(1 + |E_0|^2 + 2(E_{0,1})^2)^{-1} E_{0,1} \partial_1 s = \Lambda(t, x) \rho + \sum_{i \in \{0, 2, 3\}} \Lambda_i(t, x) \partial_i \rho + N(t, x, \varepsilon, \tilde{\rho}, \partial_1 \tilde{\rho}, \partial_2 \tilde{\rho}, \partial_3 \tilde{\rho}), \quad (3.27)$$

where $\Lambda(t, x)$ and the $\Lambda_i(t, x)$ are linear operators, and $N = (1 + |E_0|^2 + 2(E_{0,1})^2)^{-1} \tilde{N}$. By (3.18) we have

$$\|N\|_{\mathcal{H}^2(\Omega_{T_\varepsilon})} \leq C. \quad (3.28)$$

3. Let us consider the last equation in (3.12):

$$\partial_t s + \frac{1}{\varepsilon} s = -\partial_t (E_0 \cdot R + s_1 + \frac{\varepsilon}{2} |R|^2).$$

We differentiate this equation with respect to x_1 :

$$\partial_t \partial_1 s + \frac{1}{\varepsilon} \partial_1 s = -\partial_t (E_{0,1} \partial_1 R_1) - \partial_t (E_{0,2} \partial_1 R_2 + E_{0,3} \partial_1 R_3 + \partial_1 E_0 \cdot R + \partial_1 s_1 + \frac{\varepsilon}{2} \partial_1 |R|^2),$$

and using (3.27), we obtain

$$\partial_t (h \partial_1 s + b) + \frac{1}{\varepsilon} \partial_1 s = 0, \quad (3.29)$$

where

$$\begin{aligned} h &= (1 + |E_0|^2 + 2(E_{0,1})^2)^{-1} (1 + |E_0|^2), \\ b &= E_{0,1} (\Lambda \rho + \sum_{i \neq 1} \Lambda_i \partial_i \rho + N) + E_{0,2} \partial_1 R_2 + E_{0,3} \partial_1 R_3 + \partial_1 E_0 \cdot R + \partial_1 s_1 + \frac{\varepsilon}{2} \partial_1 |R|^2. \end{aligned}$$

We have

$$\frac{1}{3} \leq h \leq 1, \quad (3.30)$$

and, from (3.21), (3.28), (3.23) and (3.15),

$$\|b\|_{L^\infty(0, T_\varepsilon; L^2(\Omega))} \leq C. \quad (3.31)$$

We remark that $h\partial_1 s + b = 0$ at $t = 0$ and solving (3.29) :

$$\partial_1 s(t) = -\frac{b(t)}{h(t)} + \frac{1}{h(t)} \int_0^t \frac{1}{\varepsilon} \exp\left(-\int_\sigma^t \frac{1}{\varepsilon h(\tau)} d\tau\right) \frac{b(\sigma)}{h(\sigma)} d\sigma,$$

so, by (3.30) and (3.31),

$$\|\partial_1 s(t)\|_{L^\infty(0; T_\varepsilon; L^2(\Omega))} \leq C,$$

and by (3.27),

$$\|\partial_1 R_1\|_{L^\infty(0, T_\varepsilon; L^2(\Omega))} \leq C.$$

Therefore we obtain the \mathcal{H}^1 estimate:

$$\|\rho\|_{\mathcal{H}^1(\Omega_{T_\varepsilon})} \leq C. \quad (3.32)$$

In order to obtain the \mathcal{H}^2 estimate on the remainder term ρ , first we deal with the derivative $\partial_{1i}\rho$, $i \in \{0, 2, 3\}$. We differentiate (3.22), (3.24) and (3.26) with respect to ∂_i and using (3.21) and (3.25) we obtain

$$\|\partial_{1i} R_2\|_{L^\infty(0, T_\varepsilon; L^2(\Omega))} + \|\partial_{1i} R_3\|_{L^\infty(0, T_\varepsilon; L^2(\Omega))} \leq C, \quad (3.33)$$

$$\|\partial_{1i} S_1\|_{L^\infty(0, T_\varepsilon; L^2(\Omega))} + \|\partial_{1i} S_2\|_{L^\infty(0, T_\varepsilon; L^2(\Omega))} + \|\partial_{1i} S_3\|_{L^\infty(0, T_\varepsilon; L^2(\Omega))} \leq C.$$

We differentiate (3.29) with respect to ∂_i :

$$\partial_t(h\partial_{1,i}s + (\partial_i h\partial_1 s + \partial_i b)) + \frac{1}{\varepsilon}\partial_{1i}s = 0,$$

from (3.21), (3.33), (3.15) we remark that

$$\|\partial_i h\partial_1 s + \partial_i b\|_{L^\infty(0, T_\varepsilon; L^2(\Omega))} \leq C,$$

so, in the previous process, we can replace b by $\partial_i h\partial_1 s + \partial_i b$, and we obtain

$$\|\partial_{1i}s\|_{L^\infty(0, T_\varepsilon; L^2(\Omega))} \leq C. \quad (3.34)$$

We differentiate (3.27) with respect to ∂_i and we use (3.28), (3.21), (3.34) to obtain

$$\|\partial_{1i} R_1\|_{L^\infty(0, T_\varepsilon; L^2(\Omega))} \leq C. \quad (3.35)$$

Therefore, by (3.33)-(3.35) we have for $i \neq 1$,

$$\|\partial_{1i}\rho\|_{L^\infty(0, T_\varepsilon; L^2(\Omega))} \leq C. \quad (3.36)$$

Now we estimate $\partial_{11}\rho$. Differentiating (3.22), (3.26) and (3.24) with respect to ∂_1 , using (3.21), (3.36), (3.25) and (3.32) we obtain that

$$\|\partial_{11} R_2\|_{L^\infty(0, T_\varepsilon; L^2(\Omega))} + \|\partial_{11} R_3\|_{L^\infty(0, T_\varepsilon; L^2(\Omega))} \leq C,$$

$$\|\partial_{11} S_1\|_{L^\infty(0, T_\varepsilon; L^2(\Omega))} + \|\partial_{11} S_2\|_{L^\infty(0, T_\varepsilon; L^2(\Omega))} + \|\partial_{11} S_3\|_{L^\infty(0, T_\varepsilon; L^2(\Omega))} \leq C.$$

Differentiating (3.29) and (3.27) with respect to ∂_1 and using in particular (3.36) we obtain by the same method that

$$\|\partial_{11}s\|_{L^\infty(0, T_\varepsilon; L^2(\Omega))} + \|\partial_{11} R_1\|_{L^\infty(0, T_\varepsilon; L^2(\Omega))} \leq C,$$

which conclude the \mathcal{H}^2 estimate:

$$\|\rho\|_{\mathcal{H}^2(\Omega_{T_\varepsilon})} \leq C.$$

For the \mathcal{H}^3 estimate, we bound successively $\|\partial_{1ij}\rho\|_{L^\infty(0, T_\varepsilon; L^2(\Omega))}$ for $i, j \neq 1$, $\|\partial_{11i}\rho\|_{L^\infty(0, T_\varepsilon; L^2(\Omega))}$ for $i \neq 1$ and $\|\partial_{111}\rho\|_{L^\infty(0, T_\varepsilon; L^2(\Omega))}$. So we obtain that there exists a constant C such that

$$\|\rho\|_{\mathcal{H}^3(\Omega_{T_\varepsilon})} \leq C. \quad (3.37)$$

Last step of the proof of Theorem 3.

We recall that we fixed $T < T^*$, and that we defined T_ε by (3.15). So, either $T = T_\varepsilon$ or $\|\rho_\varepsilon\|_{\mathcal{H}^3(\Omega_{T_\varepsilon})} = \frac{1}{\sqrt{\varepsilon}}$, which is contradictory with (3.37) for ε small enough. So there exists $\varepsilon_0 > 0$ such that for $\varepsilon < \varepsilon_0$, $T_\varepsilon^* \geq T_\varepsilon = T$ and by (3.37), $\|\rho_\varepsilon\|_{\mathcal{H}^3(\Omega_T)} \leq C$, so we obtain (3.10) and (3.11) in Theorem 3.

Remark 5 *The same conclusions hold in the two-dimensional and the one-dimensional cases. In the 1-d case we can replace (3.11) by:*

$$\|e_\varepsilon - e_0\|_{\mathcal{H}^2(\Omega_T)} + \|h_\varepsilon - h_0\|_{\mathcal{H}^2(\Omega_T)} + \|v_\varepsilon\|_{\mathcal{H}^2(\Omega_T)} \leq C\varepsilon.$$

For the 1-d and 2-d TE cases, (3.10) is irrelevant since $T_\varepsilon^* = +\infty$ by Theorems 1 and 2.

4 Appendix: Proof of Proposition 2.

As in Section 2.3 the IBVP (2.4)-(2.7) is equivalent to the following system in the variable $V = (U, \chi) = (E, H, \chi)$:

$$\begin{cases} (i) & (1 + \chi)\partial_t E + (\partial_t \chi)E - \operatorname{curl} H = G_1, \\ (ii) & \partial_t H + \operatorname{curl} E = G_2, \\ (iii) & \partial_t \chi = \frac{1}{\varepsilon}(|E|^2 - \chi), \end{cases} \quad (4.1)$$

with zero initial data and homogeneous boundary condition:

$$V(0, x) = 0, x \in \Omega, \quad (4.2)$$

$$H \times n + a((E \times n) \times n) = 0, (t, x) \in \mathbb{R}^+ \times \Gamma. \quad (4.3)$$

We recall the divergence free conditions

$$\operatorname{div} H = \operatorname{div}((1 + \chi)E) = 0,$$

and the positiveness property

$$\chi \geq 0.$$

The proof is based on the following iteration scheme.

First, $\chi_k \geq 0$ being given, we define $U_{k+1} = (E_{k+1}, H_{k+1})$ by

$$\begin{cases} (1 + \chi_k)\partial_t E_{k+1} + (\partial_t \chi_k)E_{k+1} - \operatorname{curl} H_{k+1} = G_1, \\ \partial_t H_{k+1} + \operatorname{curl} E_{k+1} = G_2, \end{cases} \quad (4.4)$$

with zero initial data and homogeneous boundary condition

$$U_{k+1}(0, x) = 0, x \in \Omega, \quad (4.5)$$

$$H_{k+1} \times n + a((E_{k+1} \times n) \times n) = 0, (t, x) \in \mathbb{R}^+ \times \Gamma, \quad (4.6)$$

so the divergence free conditions hold

$$\operatorname{div} H_{k+1} = \operatorname{div}((1 + \chi_k)E_{k+1}) = 0. \quad (4.7)$$

Afterwards we define χ_{k+1} solving the differential equation

$$\partial_t \chi_{k+1} + \frac{1}{\varepsilon} \chi_{k+1} = \frac{1}{\varepsilon} |E_{k+1}|^2, \quad (4.8)$$

with the initial condition

$$\chi_{k+1}(0, x) = 0, x \in \Omega. \quad (4.9)$$

We remark that

$$\chi_{k+1} \geq 0.$$

We initialize the iteration scheme by taking $V_0 = 0$. Assuming that for any $T > 0$, $V_k \in \mathcal{H}^3(\Omega_T)$ with $\partial_t \chi_k \in \mathcal{H}^3(\Omega_T)$, using the existence results for linear problems in [16], the system (4.4)-(4.5)-(4.6) admits a unique solution $U_{k+1} = (E_{k+1}, H_{k+1}) \in \mathcal{H}^3(\Omega_T)$ which satisfies the divergence free condition (4.7). So (4.8)-(4.9) admits a unique solution $\chi_{k+1} \in \mathcal{H}^3(\Omega_T)$ and by these equations $\partial_t \chi_{k+1} \in \mathcal{H}^3(\Omega_T)$ and $\chi_{k+1} \geq 0$.

4.1 High-norm boundedness.

We denote

$$\xi_k(t) = \sum_{i=0}^3 \|\partial_t^i U_k(t)\|_{H^{3-i}(\Omega)}, \quad (4.10)$$

and

$$\varphi_k(t) = \sum_{i=0}^3 \|\partial_t^i V_k(t)\|_{H^{3-i}(\Omega)} + \sum_{i=0}^3 \|\partial_t^{i+1} \chi_k(t)\|_{H^{3-i}(\Omega)}. \quad (4.11)$$

For $\alpha = (\alpha_0, \alpha_2, \alpha_3) \in \mathbb{N}^3$ we define the operator $\partial^\alpha = \partial_t^{\alpha_0} \partial_2^{\alpha_2} \partial_3^{\alpha_3}$ and we denote

$$\eta_k(t) = \sum_{\alpha=(\alpha_0, \alpha_2, \alpha_3), |\alpha| \leq 3} \|\partial^\alpha U_k(t)\|_{L^2(\Omega)}, \quad (4.12)$$

that is, $\eta_k(t)$ measures the L^2 -norm of the tangential derivatives of U_k . Finally the source terms are estimated by

$$\Gamma(t) = \|G\|_{\mathcal{H}^3(\Omega_t)}. \quad (4.13)$$

We introduce the time T_k defined by

$$T_k = \max \left\{ t \geq 0, \sup_{s \in [0, t]} \varphi_k(s) \leq 1, \Gamma(t) \leq 1 \right\}. \quad (4.14)$$

To obtain the boundedness, we shall prove that there exists $\tilde{T} > 0$ such that

$$\forall k \geq 0, T_k \geq \tilde{T}. \quad (4.15)$$

In order to estimate the tangential derivatives of U_k we first establish the following result.

Lemma 4 *There exists a constant $K > 0$ such that, for all k , we have*

$$\forall t \leq T_k, \frac{d}{dt} \int_{\Omega} \sum_{\alpha=(\alpha_0, \alpha_2, \alpha_3), |\alpha| \leq 3} ((1 + \chi_k) |\partial^\alpha E_{k+1}|^2 + |\partial^\alpha H_{k+1}|^2) dx \leq K(1 + (\xi_{k+1})^2). \quad (4.16)$$

Proof. Taking the inner product of (4.4) with U_{k+1} we obtain

$$\begin{aligned} & \frac{1}{2} \frac{d}{dt} \int_{\Omega} ((1 + \chi_k) |E_{k+1}|^2 + |H_{k+1}|^2) dx - \int_{\partial\Omega} a((E_{k+1} \times n) \times n) \cdot E_{k+1} dx_2 dx_3 \\ & = -\frac{1}{2} \int_{\Omega} \partial_t \chi_k |E_{k+1}|^2 dx + \int_{\Omega} G \cdot U_{k+1} dx. \end{aligned}$$

Since $\|\partial_t \chi_k\|_{L^\infty(\Omega)} \leq \varphi_k$ we get the L^2 -estimate

$$\forall t \leq T_k, \frac{1}{2} \frac{d}{dt} \int_{\Omega} ((1 + \chi_k) |E_{k+1}|^2 + |H_{k+1}|^2) dx \leq K(1 + (\xi_{k+1})^2).$$

By the same method we obtain similar estimates concerning the tangential derivatives. Let us describe the order 3 derivatives estimates: for $\alpha = (\alpha_0, \alpha_2, \alpha_3)$ with $|\alpha| = \alpha_0 + \alpha_2 + \alpha_3 = 3$, we take the inner product of $\partial^\alpha(4.4)$ with $\partial^\alpha U_{k+1}$. We obtain

$$\frac{1}{2} \frac{d}{dt} \int_{\Omega} ((1 + \chi_k) |\partial^\alpha E_{k+1}|^2 + |\partial^\alpha H_{k+1}|^2) dx \leq I_1 + \dots + I_5,$$

with

$$\begin{aligned} I_1 &= K \sum_{\substack{|\beta| = |(\beta_0, \beta_2, \beta_3)| = 1 \\ |\gamma| = |(\gamma_0, \gamma_2, \gamma_3)| = 3}} \int_{\Omega} |\partial^\beta \chi_k| |\partial^\gamma E_{k+1}| |\partial^\alpha E_{k+1}| dx \\ &\leq K \sum_{|\beta|=1, |\gamma|=3} \|\partial^\beta \chi_k\|_{L^\infty(\Omega)} \|\partial^\gamma E_{k+1}\|_{L^2(\Omega)} \|\partial^\alpha E_{k+1}\|_{L^2(\Omega)}, \\ I_2 &= K \sum_{|\beta|=2, |\gamma|=2} \int_{\Omega} |\partial^\beta \chi_k| |\partial^\gamma E_{k+1}| |\partial^\alpha E_{k+1}| dx \\ &\leq K \sum_{|\beta|=2, |\gamma|=2} \|\partial^\beta \chi_k\|_{L^4(\Omega)} \|\partial^\gamma E_{k+1}\|_{L^4(\Omega)} \|\partial^\alpha E_{k+1}\|_{L^2(\Omega)}, \\ I_3 &= K \sum_{|\beta|=3, |\gamma|=1} \int_{\Omega} |\partial^\beta \chi_k| |\partial^\gamma E_{k+1}| |\partial^\alpha E_{k+1}| dx \\ &\leq K \sum_{|\beta|=3, |\gamma|=1} \|\partial^\beta \chi_k\|_{L^2(\Omega)} \|\partial^\gamma E_{k+1}\|_{L^\infty(\Omega)} \|\partial^\alpha E_{k+1}\|_{L^2(\Omega)}, \\ I_4 &= \int_{\Omega} |\partial_t \partial^\alpha \chi_k| |E_{k+1}| |\partial^\alpha E_{k+1}| dx \\ &\leq K \|\partial_t \partial^\alpha \chi_k\|_{L^2(\Omega)} \|E_{k+1}\|_{L^\infty(\Omega)} \|\partial^\alpha E_{k+1}\|_{L^2(\Omega)}, \\ I_5 &= \int_{\Omega} |\partial^\alpha G| |\partial^\alpha U_{k+1}| dx \\ &\leq \|\partial^\alpha G\|_{L^2(\Omega)} \|\partial^\alpha U_{k+1}\|_{L^2(\Omega)}. \end{aligned}$$

So using Sobolev inequalities we obtain that for all $t \leq T_k$,

$$I_1 + \dots + I_5 \leq K(1 + (\xi_{k+1})^2).$$

Now we control all the derivatives of U_{k+1} by the tangential derivatives with the following estimate.

Lemma 5 *There exists $K > 0$ such that*

$$\forall k, \forall t \leq T_k, \xi_{k+1}(t) \leq K(\eta_{k+1}(t) + \Gamma(t)). \quad (4.17)$$

Proof. In order to estimate $\partial_1 U_{k+1}$ we rewrite (4.4) and (4.7) to obtain

$$\begin{aligned}
\partial_1 H_{k+1,3} &= -\partial_t \chi_k E_{k+1,2} - (1 + \chi_k) \partial_t E_{k+1,2} + \partial_3 H_{k+1,1} + G_{1,2}, \\
\partial_1 H_{k+1,2} &= +\partial_t \chi_k E_{k+1,3} + (1 + \chi_k) \partial_t E_{k+1,3} + \partial_2 H_{k+1,1} - G_{1,3}, \\
\partial_1 E_{k+1,3} &= \partial_3 E_{k+1,1} + \partial_t H_{k+1,2} - G_{2,2}, \\
\partial_1 E_{k+1,2} &= \partial_2 E_{k+1,1} - \partial_t H_{k+1,3} + G_{2,3}, \\
\partial_1 H_{k+1,1} &= -\partial_2 H_{k+1,2} - \partial_3 H_{k+1,3}, \\
\partial_1 E_{k+1,1} &= -\partial_2 E_{k+1,2} - \partial_3 E_{k+1,3} - (1 + \chi_k)^{-1} (\nabla \chi_k \cdot E_{k+1}).
\end{aligned} \tag{4.18}$$

Using (4.10), (4.11), (4.12), (4.13) and (4.14), we obtain the estimate

$$\forall t \leq T_k, \|\partial_1 U_{k+1}\|_{L^2(\Omega)} \leq K(\eta_{k+1}(t) + \Gamma(t)).$$

In order to estimate $\partial_{1i} U_{k+1}$ we differentiate (4.18) with respect to ∂_i . For $i \neq 1$, we obtain directly that

$$\forall t \leq T_k, \|\partial_{1i} U_{k+1}\|_{L^2(\Omega)} \leq K(\eta_{k+1}(t) + \Gamma(t)), \tag{4.19}$$

and using (4.19) we obtain

$$\forall t \leq T_k, \|\partial_{11} U_{k+1}\|_{L^2(\Omega)} \leq K(\eta_{k+1}(t) + \Gamma(t)).$$

With the same arguments we estimate successively $\|\partial_{1ij} U_{k+1}\|_{L^2(\Omega)}$, $\|\partial_{11j} U_{k+1}\|_{L^2(\Omega)}$ and $\|\partial_{111} U_{k+1}\|_{L^2(\Omega)}$ where $i \neq 1$ and $j \neq 1$, so we conclude the proof of Lemma 5.

Using (4.16) and (4.17), by Gronwall lemma, we obtain that there exists $K > 0$ such that

$$\forall k, \forall t \leq T_k, (\eta_{k+1}(t))^2 \leq e^{Kt} - 1,$$

so by (4.17), there exists $K > 0$ such that

$$\forall k, \forall t \leq T_k, \xi_{k+1}(t) \leq K \left((e^{Kt} - 1)^{\frac{1}{2}} + \Gamma(t) \right). \tag{4.20}$$

Then, solving (4.8) we obtain that

$$\|\chi_{k+1}(t)\|_{H^3(\Omega)} \leq K \sup_{s \in [0,t]} (\xi_{k+1}(s))^2.$$

Using (4.8), $\partial_t(4.8)$, $\partial_t^2(4.8)$ and $\partial_t^3(4.8)$ we obtain that

$$\sum_{i=1}^4 \|\partial_t^i \chi_{k+1}\|_{H^{4-i}(\Omega)} \leq K \sup_{s \in [0,t]} (\xi_{k+1}(s))^2.$$

So there exists $K > 0$ such that

$$\forall k, \varphi_{k+1}(t) \leq \xi_{k+1}(t) + K \sup_{s \in [0,t]} (\xi_{k+1}(s))^2.$$

Therefore with (4.20) there exists $K > 0$ such that

$$\forall k, \forall t \leq T_k, \varphi_{k+1}(t) \leq g(t) := K \left((e^{Kt} - 1)^{\frac{1}{2}} + \Gamma(t) + ((e^{Kt} - 1)^{\frac{1}{2}} + \Gamma(t))^2 \right).$$

Since $g(0) = 0$ there exists $\tilde{T} > 0$ such that for $t \leq \tilde{T}$, $g(t) \leq \frac{3}{4}$ and $\Gamma(t) \leq \frac{3}{4}$, so we conclude the proof of (4.15). In addition we have obtained the high-norm boundedness:

$$\forall k, \sup_{[0,\tilde{T}]} \varphi_k(t) \leq 1. \tag{4.21}$$

4.2 Low-norm contraction.

Subtracting (4.4) for $k-1$ to the one for k we have

$$\begin{cases} (1 + \chi_k)\partial_t(E_{k+1} - E_k) - \operatorname{curl}(H_{k+1} - H_k) &= -\partial_t\chi_k(E_{k+1} - E_k) + (\chi_{k-1} - \chi_k)\partial_tE_k \\ &+ (\partial_t\chi_{k-1} - \partial_t\chi_k)E_k, \\ \partial_t(H_{k+1} - H_k) + \operatorname{curl}(E_{k+1} - E_k) &= 0, \end{cases} \quad (4.22)$$

with the zero initial data and the homogeneous boundary condition

$$(E_{k+1} - E_k)(0, x) = (H_{k+1} - H_k)(0, x) = 0, \quad x \in \Omega,$$

$$(H_{k+1} - H_k) \times n + a(((E_{k+1} - E_k) \times n) \times n) = 0, \quad (t, x) \in \mathbb{R}^+ \times \Gamma.$$

Taking the inner product of (4.22) with $U_{k+1} - U_k$ we obtain

$$\begin{aligned} \frac{1}{2} \frac{d}{dt} \int_{\Omega} ((1 + \chi_k)|E_{k+1} - E_k|^2 + |H_{k+1} - H_k|^2) dx &\leq -\frac{1}{2} \int_{\Omega} \partial_t\chi_k |E_{k+1} - E_k|^2 dx \\ + \int_{\Omega} (\chi_{k-1} - \chi_k)\partial_tE_k(E_{k+1} - E_k) dx &+ \int_{\Omega} (\partial_t\chi_{k-1} - \partial_t\chi_k)E_k \cdot (E_{k+1} - E_k) dx, \end{aligned}$$

so using (4.21)

$$\begin{aligned} &\frac{d}{dt} \int_{\Omega} ((1 + \chi_k)|E_{k+1} - E_k|^2 + |H_{k+1} - H_k|^2) dx \\ &\leq K \left(\|E_{k+1} - E_k\|_{L^2(\Omega)}^2 + \|\chi_k - \chi_{k-1}\|_{L^2(\Omega)}^2 + \|\partial_t\chi_k - \partial_t\chi_{k-1}\|_{L^2(\Omega)}^2 \right). \end{aligned} \quad (4.23)$$

Subtracting (4.8) for $k-2$ to the one for $k-1$ we have

$$\partial_t(\chi_k - \chi_{k-1}) + \frac{1}{\varepsilon}(\chi_k - \chi_{k-1}) = \frac{1}{\varepsilon}(E_k + E_{k-1}) \cdot (E_k - E_{k-1}), \quad (4.24)$$

so

$$\chi_k - \chi_{k-1} = \int_0^t \frac{1}{\varepsilon} \exp\left(-\frac{t-s}{\varepsilon}\right) (E_k + E_{k-1})(s) \cdot (E_k - E_{k-1})(s) ds. \quad (4.25)$$

We introduce

$$u_k(t) = \|U_{k+1} - U_k\|_{L^\infty(0,t;L^2(\Omega))},$$

and from (4.25) and (4.21) we have

$$\|(\chi_k - \chi_{k-1})(t)\|_{L^2(\Omega)} \leq K u_{k-1}(t), \quad (4.26)$$

and from (4.24)

$$\|(\partial_t\chi_k - \partial_t\chi_{k-1})(t)\|_{L^2(\Omega)} \leq K u_{k-1}(t). \quad (4.27)$$

Integrating (4.23) and using (4.26) and (4.27) we have

$$(u_k(t))^2 \leq K t (u_k(t))^2 + K t (u_{k-1}(t))^2. \quad (4.28)$$

We fix $\bar{T} \leq \tilde{T}$ such that, in (4.28), $K\bar{T} \leq \frac{1}{4}$. So we have

$$\forall k, \forall t \leq \bar{T}, u_k(t) \leq \frac{1}{\sqrt{3}} u_{k-1}(t).$$

Hence U_k is a Cauchy sequence in $\mathcal{C}(0, \bar{T}; L^2(\Omega))$ and by (4.26), χ_k is also a Cauchy sequence in $\mathcal{C}(0, \bar{T}; L^2(\Omega))$. By standard arguments we obtain a smooth local solution to (4.1)-(4.2)-(4.3). The end of the proof of Proposition 2 is classical.

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