

# Numerical schemes for complex PDEs & MSCA-IF: SuPerMan

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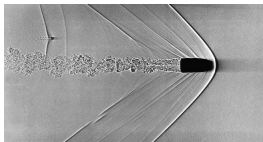
21/04/2021

## New robust and effective numerical methods for geophysical and astrophysical flows

- Vortical flows: new energies, astrophysics



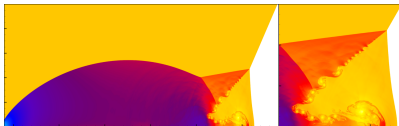
- Multiphase flows: instabilities, fluid-structure interaction



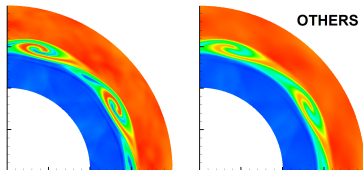
**Common mathematical description**  
**Hyperbolic PDEs:**  $\partial_t \mathbf{Q} + \nabla \cdot \mathbf{F}(\mathbf{Q}) + \mathbf{B}(\mathbf{Q}) \cdot \nabla \mathbf{Q} = \mathbf{S}(\mathbf{Q})$

# Difficulties & corresponding numerical schemes capabilities

- Multi-scale phenomena



- Equilibria and physical instabilities



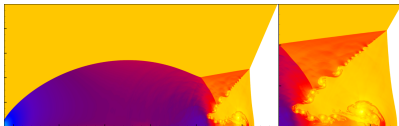
- Interface tracking

- Convection without dissipation

# Difficulties & corresponding numerical schemes capabilities

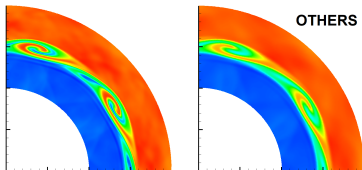
High order FV-DG methods + HPC (Fortran MPI or CUDA)

- Multi-scale phenomena



- Interface tracking

- Equilibria and physical instabilities

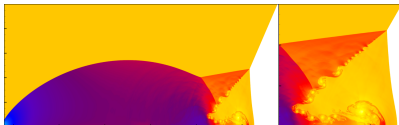


- Convection without dissipation

# Difficulties & corresponding numerical schemes capabilities

High order FV-DG methods + HPC (Fortran MPI or CUDA)

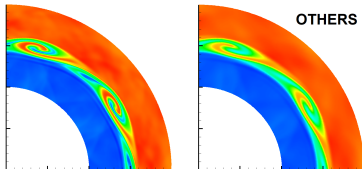
- Multi-scale phenomena



AMR3D

- Interface tracking

- Equilibria and physical instabilities



⇒ well balanced methods

Nonconforming-WB

CUDA-WB

WB-GR

- Convection without dissipation

Lagrangian schemes

ALEVoronoi

# 1. Multiphase: interfaces and fluid structure interaction

- **Diffuse interface methods:** Idea: Model + high order

Volume fraction  $\alpha$  transported with the interface velocity  $\mathbf{v}_I$   $\partial_t \alpha + \mathbf{v}_I \cdot \nabla \alpha = 0$   
coupled with the multiphase model

Rotating solids 2D

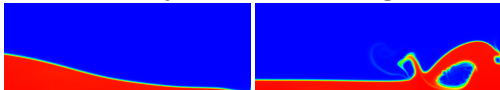
AMR3D

$$\frac{\partial(\alpha\rho)}{\partial t} + \frac{\partial(\alpha\rho v_k)}{\partial x_k} = 0$$
$$\frac{\partial(\alpha\rho v_i)}{\partial t} + \frac{\partial(\alpha\rho v_i v_k + \alpha p \delta_{ik})}{\partial x_k} = \rho g_i$$

Rotating solids 3D

CUDA-WB

Fluids: dambreak and  
non hydrostatic breaking waves



- **Sharp interface methods:** Idea: Geometry + high order

CARDAMOM, in particular: MR & M. Ciallella (PhD)

HOTHYPE

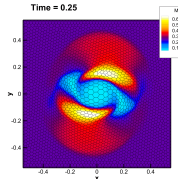
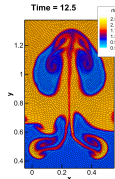
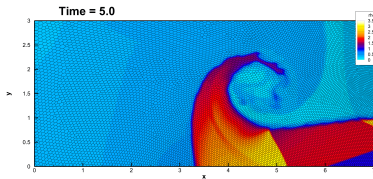
## 2. Direct ALE on moving meshes with **topology changes**

Euler: vortex on moving meshes  
with reconnection

DG $P_2P_2 \rightarrow \mathcal{O}3$		
$h_f$	$\epsilon(\rho)_{L_1}$	$\mathcal{O}$
7.5E-01	1.4E-02	-
6.1E-01	7.2E-03	3.4
3.2E-01	9.3E-04	3.2
2.2E-01	2.8E-04	3.0
1.6E-01	1.2E-04	3.0
DG $P_3P_3 \rightarrow \mathcal{O}4$		
$h_f$	$\epsilon(\rho)_{L_1}$	$\mathcal{O}$
6.1E-01	1.4E-03	-
5.2E-01	7.4E-04	3.7
4.7E-01	4.1E-04	5.9
3.2E-01	7.7E-05	4.4
2.2E-01	1.6E-05	4.0
DG $P_4P_4 \rightarrow \mathcal{O}5$		
$h_f$	$\epsilon(\rho)_{L_1}$	$\mathcal{O}$
1.4E-00	1.1e-02	-
1.0E-00	2.0e-03	5.9
9.8E-01	1.6e-03	4.7
8.9E-01	9.0e-04	5.9
8.5E-01	7.0e-04	5.1

Multiphase: triple point and Rayleigh-Taylor

MHD: rotor



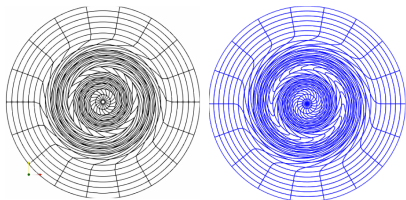
ALEVoronoi

to be improved by coupling with

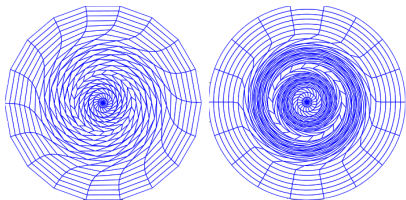
MMG

## 2. Lagrangian methods: meshes and vortical fluxes

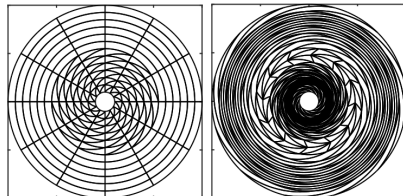
- **Shashkov & Morgan - USA**



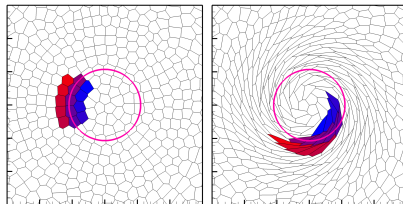
- **Maire & Vilar - CEA**



- **Standard ALE approach**



- **Voronoi meshes**

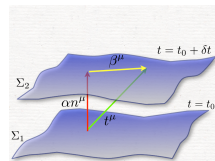




### 3. MSCA-IF funded project (184.700€)

## SuPerMan

Structure Preserving schemes for conservation laws  
on space time Manifolds



$$\partial_t \mathbf{Q} + \nabla \cdot \mathbf{F}(\mathbf{Q}) + \mathbf{B}(\mathbf{Q}) \cdot \nabla \mathbf{Q} = \mathbf{S}(\mathbf{Q})$$

$$\boxed{\mathbf{Q} \leftarrow \gamma_{ij}} \quad \text{with} \quad ds^2 = -(\alpha^2 - \beta^i \beta_i) dt^2 + 2\beta_i dx^i dt + \gamma_{ij} dx^i dx^j$$

#### Schemes

- **HPC** & **Lagrangian** of very high order
- **asymptotic preserving**, **well balanced** and **metric independent**

for applications in

- **Geophysics**: tsunami/hurricane on the geoid
- **Astrophysics**: vortical fluxes around black holes/neutron stars

#### National/international network:

INRIA (MR + 1 postdoc from 2022) + Uni Trento + Max Planck + Uni Malaga

→ fundation for **ERC StG** application

### 3. CCZ4: Einstein field equations (hyp. order 1, 59 éq.)

$$\begin{aligned}
 \partial_t \tilde{\gamma}_{ij} &= \beta^k 2D_{kij} + \tilde{\gamma}_{ki} B_j^k + \tilde{\gamma}_{kj} B_i^k - 2/3 \tilde{\gamma}_{ij} B_k^k - 2\alpha \left( \tilde{A}_{ij} - 1/3 \tilde{\gamma}_{ij} \text{tr} \tilde{A} \right) - \tau^{-1} (\tilde{\gamma} - 1) \tilde{\gamma}_{ij}, \\
 \partial_t \ln \alpha &= \beta^k A_k - \alpha g(\alpha) (K - K_0 - 2\Theta c), \\
 \partial_t \beta^i &= s \beta^k B_k^i + s f b^i \\
 \partial_t \ln \phi &= \beta^k P_k + 1/3 \left( \alpha K - B_k^k \right), \\
 \partial_t \tilde{A}_{ij} - \beta^k \partial_k \tilde{A}_{ij} &= \phi^2 \left[ -\nabla_i \nabla_j \alpha + \alpha \left( R_{ij} + \nabla_i Z_j + \nabla_j Z_i \right) \right] + \phi^2 1/3 \frac{\tilde{\gamma}_{ij}}{\phi^2} \left[ -\nabla^k \nabla_k \alpha + \alpha (R + 2\nabla_k Z^k) \right] \\
 &= \tilde{A}_{ki} B_j^k + \tilde{A}_{kj} B_i^k - 2/3 \tilde{A}_{ij} B_k^k + \alpha \tilde{A}_{ij} (K - 2\Theta c) - 2\alpha \tilde{A}_{ij} \tilde{\gamma}^{lm} \tilde{A}_{mj} - \tau^{-1} \tilde{\gamma}_{ij} \text{tr} \tilde{A}, \\
 \partial_t K - \beta^k \partial_k K &+ \nabla^i \nabla_i \alpha - \alpha (R + 2\nabla_i Z^i) = \alpha K (K - 2\Theta c) - 3\alpha \kappa_1 (1 + \kappa_2) \Theta \\
 \partial_t \Theta - \beta^k \partial_k \Theta &- 1/2 \alpha \epsilon^2 (R + 2\nabla_i Z^i) = 1/2 \alpha \epsilon^2 \left( 2/3 K^2 - \tilde{A}_{ij} \tilde{A}^{ij} \right) - \alpha \Theta K c - Z^i \alpha A_i - \alpha \kappa_1 (2 + \kappa_2) \Theta, \\
 \partial_t \hat{\Gamma}^i - \beta^k \partial_k \hat{\Gamma}^i &+ 4/3 \alpha \tilde{\gamma}^{ij} \partial_j K - 2\alpha \tilde{\gamma}^{ki} \partial_k \Theta - s \tilde{\gamma}^{kl} \partial_{(k} B_{l)}^i - s 1/3 \tilde{\gamma}^{ik} \partial_{(k} B_{l)}^j - s 2\alpha \tilde{\gamma}^{ik} \tilde{\gamma}^{nm} \partial_k \tilde{A}_{nm} \\
 &= 2/3 \tilde{\Gamma}^i B_k^k - \tilde{\Gamma}^k B_k^i + 2\alpha \left( \tilde{\Gamma}^i_{jk} \tilde{A}^{jk} - 3 \tilde{A}^{ij} P_j \right) - 2\alpha \tilde{\gamma}^{ki} (\Theta A_k + 2/3 K Z_k) - 2\alpha \tilde{A}^{ij} A_j \\
 &\quad - 4s \alpha \tilde{\gamma}^{ik} D_k^{nm} \tilde{A}_{nm} + 2\kappa_3 \left( 2/3 \tilde{\gamma}^{ij} Z_j B_k^k - \tilde{\gamma}^{jk} Z_j B_k^i \right) - 2\alpha \kappa_1 \tilde{\gamma}^{ij} Z_j \\
 \partial_t b^i - s \beta^k \partial_k b^i &= s \left( \partial_t \hat{\Gamma}^i - \beta^k \partial_k \hat{\Gamma}^i - \eta b^i \right),
 \end{aligned}$$

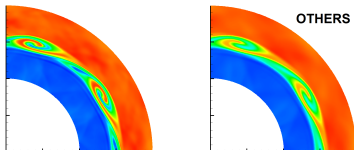
avec les EDP suivantes pour les variables auxiliaires

$$\begin{aligned}
 \partial_t A_k - \beta^l \partial_l A_k &+ \alpha g(\alpha) (\partial_k K - \partial_k K_0 - 2c \partial_k \Theta) + s \alpha g(\alpha) \tilde{\gamma}^{nm} \partial_k \tilde{A}_{nm} \\
 &= +2s \alpha g(\alpha) D_k^{nm} \tilde{A}_{nm} - \alpha A_k (K - K_0 - 2\Theta c) \left( g(\alpha) + \alpha g'(\alpha) \right) + B_k^l A_l, \\
 \partial_t B_k^i - s \beta^l \partial_l B_k^i &- s \left( f \partial_k b^i + \alpha^2 \mu \tilde{\gamma}^{ij} \left( \partial_k P_j - \partial_j P_k \right) - \alpha^2 \mu \tilde{\gamma}^{ij} \tilde{\gamma}^{nl} \left( \partial_k D_{ljn} - \partial_l D_{kjn} \right) \right) = s B_k^l B_l^i, \\
 \partial_t D_{kij} - \beta^l \partial_l D_{kij} &+ s \left( -1/2 \tilde{\gamma}^{mi} \partial_{(k} B_{j)}^m - 1/2 \tilde{\gamma}^{mj} \partial_{(k} B_{i)}^m + 1/3 \tilde{\gamma}_{ij} \partial_{(k} B_{m)}^m \right) + \alpha \partial_k \tilde{A}_{ij} - \alpha 1/3 \tilde{\gamma}_{ij} \tilde{\gamma}^{nm} \partial_k \tilde{A}_{nm} \\
 &= B_k^l D_{lij} + B_j^l D_{kli} + B_i^l D_{klj} - 2/3 B_l^j D_{kij} - \alpha 2/3 \tilde{\gamma}_{ij} D_k^{nm} \tilde{A}_{nm} - \alpha A_k \left( \tilde{A}_{ij} - 1/3 \tilde{\gamma}_{ij} \text{tr} \tilde{A} \right), \\
 \partial_t P_k - \beta^l \partial_l P_k &- 1/3 \alpha \partial_k K + s 1/3 \partial_{(k} B_{i)}^j - s 1/3 \alpha \tilde{\gamma}^{nm} \partial_k \tilde{A}_{nm} \\
 &= 1/3 \alpha A_k K + B_k^l P_l - s 2/3 \alpha D_k^{nm} \tilde{A}_{nm}.
 \end{aligned}$$

# 3. Preliminary results in astrophysics

- Well balancing for Euler equations (2D)

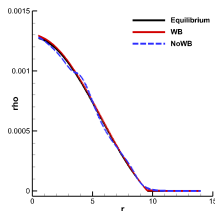
Nonconforming-WB



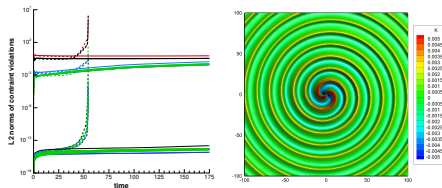
- Well balancing for General Relativistic Magnhydrodynamics

1D (GRMHD)

WB-GR



- GLM curl cleaning for Einstein field equations 3D (CCZ4)



**Merci beaucoup pour votre attention !**